

Precise predictions for radiative return at flavour factories

14th edition of the International Workshop on e^+e^-
collisions from Phi to Psi
(PhiPsi 2026)

Marco Ghilardi
On behalf of the BabaYaga team

June 9, 2026



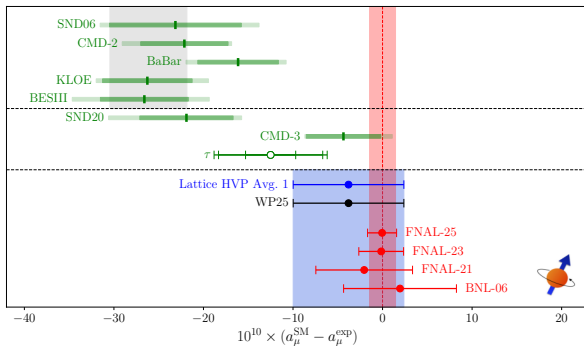
UNIVERSITÀ
DI PAVIA

The Muon anomaly

$$\vec{\mu} = -g\mu_B\vec{S} \quad a_\mu = \frac{(g-2)_\mu}{2}$$

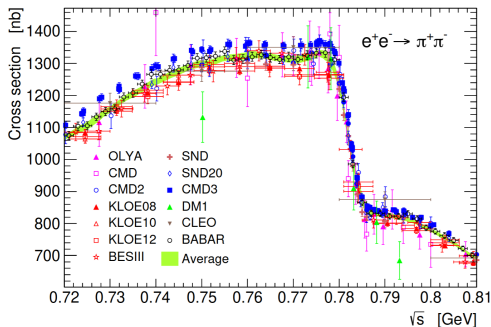
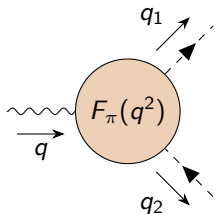
Muon $g-2$

The current status of the theory–experiment comparison is the following:



The Pion Form Factor

$$\langle \pi^+(q_2) \pi^-(q_1) | J_\pi^\mu(0) | 0 \rangle = e(q_1 - q_2)^\mu F_\pi(q^2)$$



The radiative return



At fixed collider energy \sqrt{s} , lower-energy two pion pairs can be studied via initial-state radiation:

$$e^+e^- \rightarrow \pi^+\pi^-\gamma, \quad M_{\pi\pi}^2 = (p_{\pi^+} + p_{\pi^-})^2 < s$$

The radiative return



At fixed collider energy \sqrt{s} , lower-energy two pion pairs can be studied via initial-state radiation:

$$e^+e^- \rightarrow \pi^+\pi^-\gamma, \quad M_{\pi\pi}^2 = (p_{\pi^+} + p_{\pi^-})^2 < s$$

The measured differential cross section relates to the pion production cross section:

$$\frac{d\sigma_{\pi^+\pi^-\gamma}}{dM_{\pi\pi}^2} = H(s, M_{\pi\pi}^2) \sigma_{\pi^+\pi^-}(M_{\pi\pi}^2),$$

where $H(s, M_{\pi\pi}^2)$ is the radiator function.

The radiative return



At fixed collider energy \sqrt{s} , lower-energy two pion pairs can be studied via initial-state radiation:

$$e^+e^- \rightarrow \pi^+\pi^-\gamma, \quad M_{\pi\pi}^2 = (p_{\pi^+} + p_{\pi^-})^2 < s$$

The measured differential cross section relates to the pion production cross section:

$$\frac{d\sigma_{\pi^+\pi^-\gamma}}{dM_{\pi\pi}^2} = H(s, M_{\pi\pi}^2) \sigma_{\pi^+\pi^-}(M_{\pi\pi}^2),$$

where $H(s, M_{\pi\pi}^2)$ is the radiator function.

A precise theoretical treatment of radiative corrections is required to extract the pion form factor at the sub-percent level.

The perturbative accuracy

LL resummation

Approximate calculation of soft-virtual corrections to any order.

$$1 \quad \text{---}$$

$$c_{\alpha}^{LL} \quad \text{---} \text{---} \text{---} \text{---} \text{---}$$

$$\frac{(c_{\alpha}^{LL})^2}{2} \quad \text{---} \text{---} \text{---} \text{---} \text{---} \text{---} \text{---}$$

NLO

Exact calculation up to $\mathcal{O}(\alpha)$ w.r.t. the born scattering amplitude.

- Virtual corrections



- Real corrections



The perturbative accuracy

LL resummation

Approximate calculation of soft-virtual corrections to any order.

$$1 \quad \text{---} \rightarrow \text{---}$$

$$c_{\alpha}^{LL} \quad \text{---} \overset{\text{loop}}{\curvearrowright} \text{---} \quad \text{---} \overset{\text{collinear}}{\curvearrowright} \text{---}$$

$$\frac{(c_{\alpha}^{LL})^2}{2} \quad \text{---} \overset{\text{collinear}}{\curvearrowright} \text{---} \quad \text{---} \overset{\text{loop}}{\curvearrowright} \text{---} \overset{\text{collinear}}{\curvearrowright} \text{---} \quad \text{---} \overset{\text{loop}}{\curvearrowright} \text{---} \overset{\text{loop}}{\curvearrowright} \text{---}$$

NLO

Exact calculation up to $\mathcal{O}(\alpha)$ w.r.t. the born scattering amplitude.

- Virtual corrections



- Real corrections



The perturbative accuracy

LL resummation

Approximate calculation of soft-virtual corrections to any order.

$$\begin{array}{lll}
 \alpha^0 & & \\
 \alpha L & \alpha & \\
 \frac{1}{2}\alpha^2 L^2 & \frac{1}{2}\alpha^2 L & \frac{1}{2}\alpha^2 \\
 \sum_{n=3}^{\infty} \frac{\alpha^n}{n!} L^n & \sum_{n=3}^{\infty} \frac{\alpha^n}{n!} L^{n-1} & \dots
 \end{array}$$

NLO

Exact calculation up to $\mathcal{O}(\alpha)$ w.r.t. the born scattering amplitude.

$$\begin{array}{lll}
 \alpha^0 & & \\
 \alpha L & \alpha & \\
 \frac{1}{2}\alpha^2 L^2 & \frac{1}{2}\alpha^2 L & \frac{1}{2}\alpha^2 \\
 \sum_{n=3}^{\infty} \frac{\alpha^n}{n!} L^n & \sum_{n=3}^{\infty} \frac{\alpha^n}{n!} L^{n-1} & \dots
 \end{array}$$

The perturbative accuracy



NLO+LL (NLOPS)

LO	α^0		
NLO	αL	α	
NNLO	$\frac{1}{2}\alpha^2 L^2$	$\frac{1}{2}\alpha^2 L$	$\frac{1}{2}\alpha^2$
h.o.	$\sum_{n=3}^{\infty} \frac{\alpha^n}{n!} L^n$	$\sum_{n=3}^{\infty} \frac{\alpha^n}{n!} L^{n-1}$...

The perturbative accuracy

NLO+LL (NLOPS)

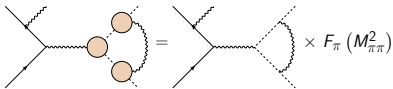
LO	α^0		
NLO	αL	α	
NNLO	$\frac{1}{2}\alpha^2 L^2$	$\frac{1}{2}\alpha^2 L$	$\frac{1}{2}\alpha^2$
h.o.	$\sum_{n=3}^{\infty} \frac{\alpha^n}{n!} L^n$	$\sum_{n=3}^{\infty} \frac{\alpha^n}{n!} L^{n-1}$...

NLOPS master formula

$$d\sigma_{\text{NLOPS}} = \sum_{n=1}^{\infty} \exp\{-C_{\alpha}^{\text{LL}}\} F_{\text{SV}} \frac{1}{n!} \left| \mathcal{M}_n^{\text{X}} \right|^2 d\Phi_n(\{p\}, \{k\}).$$

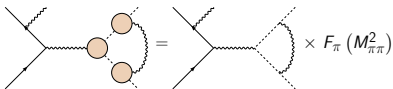
The Pion internal structure

The standard way which respects IR cancellation is $F \times \text{sQED}$:



The Pion internal structure

The standard way which respects IR cancellation is $F \times s\text{QED}$:



$$= \times F_{\pi}(M_{\pi\pi}^2)$$

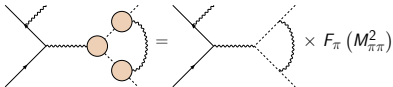
 $F \times s\text{QED}$

2601.19530



The Pion internal structure

The standard way which respects IR cancellation is $F \times s$ QED:

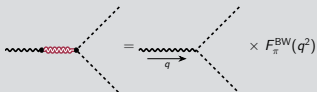


To include the pion form factor into loop integration:

GVMD

$$F_\pi^{\text{BW}}(q^2) = \sum_{v=1}^{n_r} F_{\pi,v}^{\text{BW}}(q^2)$$

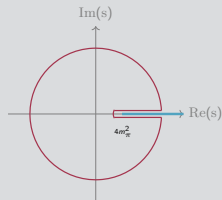
$$F_{\pi,v}^{\text{BW}}(q^2) = \frac{c_v}{c_t} \frac{\Lambda_v^2}{\Lambda_v^2 - q^2}$$



FsQED

$$F_\pi(q^2) = 1 - \frac{q^2}{\pi} \int_{4m_\pi^2}^{\infty} \frac{ds'}{s'} \frac{\text{Im}F_\pi(s')}{q^2 - s'}$$

$$\frac{1}{\pi} \int_{4m_\pi^2}^{\infty} \frac{ds'}{s'} \text{Im}F_\pi(s') = 1$$



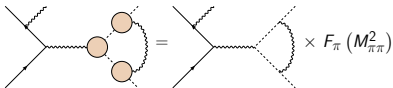
$F \times s$ QED

2601.19530



The Pion internal structure

The standard way which respects IR cancellation is $F \times s$ QED:

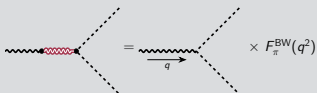


To include the pion form factor into loop integration:

GVMD

$$F_\pi^{\text{BW}}(q^2) = \sum_{v=1}^{n_r} F_{\pi,v}^{\text{BW}}(q^2)$$

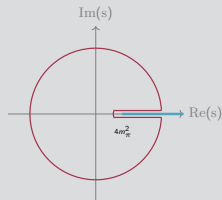
$$F_{\pi,v}^{\text{BW}}(q^2) = \frac{c_v}{c_t} \frac{\Lambda_v^2}{\Lambda_v^2 - q^2}$$



FsQED

$$F_\pi(q^2) = 1 - \frac{q^2}{\pi} \int_{4m_\pi^2}^{\infty} \frac{ds'}{s'} \frac{\text{Im}F_\pi(s')}{q^2 - s'}$$

$$\frac{1}{\pi} \int_{4m_\pi^2}^{\infty} \frac{ds'}{s'} \text{Im}F_\pi(s') = 1$$



F × sQED

2601.19530



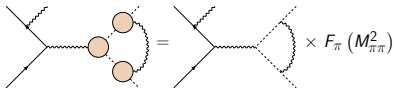
GVMD

2603.28621



The Pion internal structure

The standard way which respects IR cancellation is $F \times s$ QED:

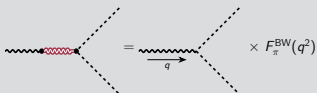


To include the pion form factor into loop integration:

GVMD

$$F_\pi^{\text{BW}}(q^2) = \sum_{v=1}^{n_r} F_{\pi,v}^{\text{BW}}(q^2)$$

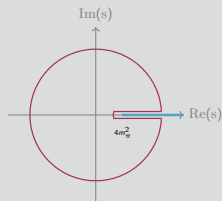
$$F_{\pi,v}^{\text{BW}}(q^2) = \frac{c_v}{c_t} \frac{\Lambda_v^2}{\Lambda_v^2 - q^2}$$



FsQED

$$F_\pi(q^2) = 1 - \frac{q^2}{\pi} \int_{4m_\pi^2}^{\infty} \frac{ds'}{s'} \frac{\text{Im}F_\pi(s')}{q^2 - s'}$$

$$\frac{1}{\pi} \int_{4m_\pi^2}^{\infty} \frac{ds'}{s'} \text{Im}F_\pi(s') = 1$$



F × sQED

2601.19530



GVMD

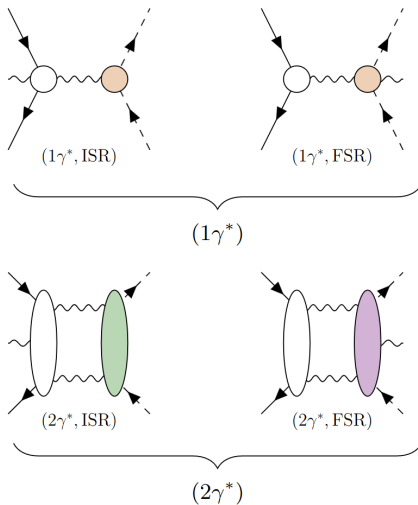
2603.28621



FsQED

W.I.P.

Subsets of GVMD corrections



<p>(a) KLOE I Large-Angle (LA)</p> <p>$\sqrt{s} = 1.02 \text{ GeV}$ $50^\circ \leq \theta^\pm \leq 130^\circ$ $p_z^\pm \geq 90 \text{ MeV} \vee p_\perp^\pm \geq 160 \text{ MeV}$ $50^\circ \leq \theta_\gamma \leq 130^\circ \wedge E_\gamma \geq 20 \text{ MeV}$ $0.1 \text{ GeV}^2 \leq M_{XX}^2 \leq 0.85 \text{ GeV}^2$</p>	<p>(b) KLOE II Small-Angle (SA)</p> <p>$\sqrt{s} = 1.02 \text{ GeV}$ $50^\circ \leq \theta^\pm \leq 130^\circ$ $p_z^\pm \geq 90 \text{ MeV} \vee p_\perp^\pm \geq 160 \text{ MeV}$ $\theta_\gamma \leq 15^\circ \vee \theta_\gamma \geq 165^\circ$ $0.35 \text{ GeV}^2 \leq M_{XX}^2 \leq 0.95 \text{ GeV}^2$</p>
<p>(c) BES III</p> <p>$\sqrt{s} = 4 \text{ GeV}$ $\cos \theta^\pm \leq 0.93 \wedge p_\perp^\pm \geq 300 \text{ MeV}$ $[\cos \theta_\gamma \leq 0.8 \wedge E_\gamma \geq 25 \text{ MeV}]$ $\vee [0.86 \leq \cos \theta_\gamma \leq 0.92 \wedge E_\gamma \geq 50 \text{ MeV}]$ $\exists! \gamma : E_\gamma \geq 400 \text{ MeV}$</p>	<p>(d) B</p> <p>$\sqrt{s} = 10 \text{ GeV}$ $0.65 \text{ rad} \leq \theta^\pm \leq 2.75 \text{ rad} \wedge p^\pm \geq 1 \text{ GeV}$ $0.6 \text{ rad} \leq \theta_\gamma \leq 2.7 \text{ rad} \wedge E_\gamma \geq 3 \text{ GeV}$ $\theta_{\bar{\gamma}, \gamma^{(h)}} \leq 0.3 \text{ rad}$ $M_{XX\gamma} \geq 8 \text{ GeV}$</p>

(a) KLOE I Large-Angle (LA)

$$\sqrt{s} = 1.02 \text{ GeV}$$

$$50^\circ \leq \theta^\pm \leq 130^\circ$$

$$p_z^\pm \geq 90 \text{ MeV} \vee p_\perp^\pm \geq 160 \text{ MeV}$$

$$50^\circ \leq \theta_\gamma \leq 130^\circ \wedge E_\gamma \geq 20 \text{ MeV}$$

$$0.1 \text{ GeV}^2 \leq M_{XX}^2 \leq 0.85 \text{ GeV}^2$$

(b) KLOE II Small-Angle (SA)

$$\sqrt{s} = 1.02 \text{ GeV}$$

$$50^\circ \leq \theta^\pm \leq 130^\circ$$

$$p_z^\pm \geq 90 \text{ MeV} \vee p_\perp^\pm \geq 160 \text{ MeV}$$

$$\theta_\gamma \leq 15^\circ \vee \theta_\gamma \geq 165^\circ$$

$$0.35 \text{ GeV}^2 \leq M_{XX}^2 \leq 0.95 \text{ GeV}^2$$

(c) BES III

$$\sqrt{s} = 4 \text{ GeV}$$

$$|\cos \theta^\pm| \leq 0.93 \wedge p_\perp^\pm \geq 300 \text{ MeV}$$

$$[|\cos \theta_\gamma| \leq 0.8 \wedge E_\gamma \geq 25 \text{ MeV}]$$

$$\vee [0.86 \leq |\cos \theta_\gamma| \leq 0.92 \wedge E_\gamma \geq 50 \text{ MeV}]$$

$$\exists! \gamma : E_\gamma \geq 400 \text{ MeV}$$

(d) B

$$\sqrt{s} = 10 \text{ GeV}$$

$$0.65 \text{ rad} \leq \theta^\pm \leq 2.75 \text{ rad} \wedge p^\pm \geq 1 \text{ GeV}$$

$$0.6 \text{ rad} \leq \theta_\gamma \leq 2.7 \text{ rad} \wedge E_\gamma \geq 3 \text{ GeV}$$

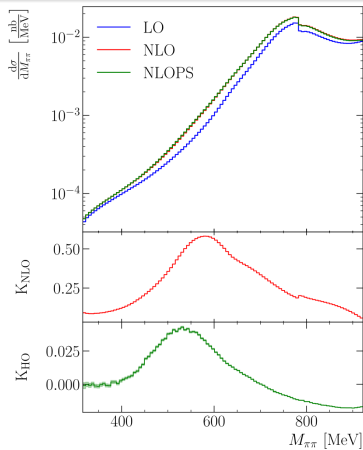
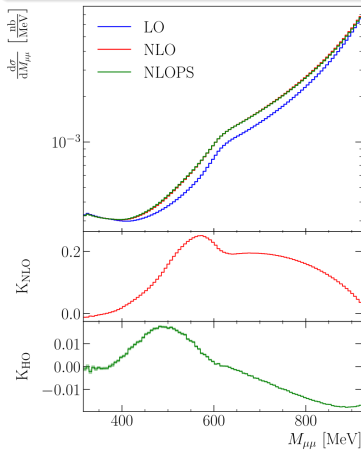
$$\theta_{\bar{\gamma}, \gamma^{(h)}} \leq 0.3 \text{ rad}$$

$$M_{XX\gamma} \geq 8 \text{ GeV}$$

$M_{\mu\mu}, M_{\pi\pi} - F \times s\text{QED}$

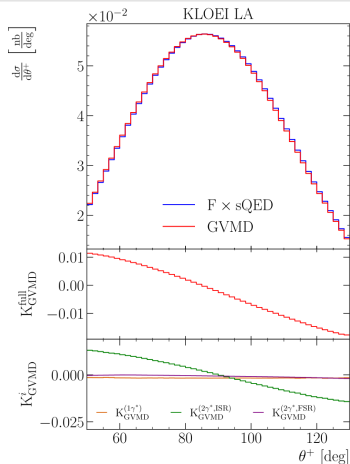
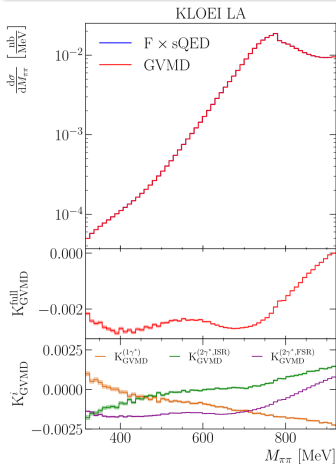
$$K_{\text{NLO}} = \frac{d\sigma_{\text{NLO}} - d\sigma_{\text{LO}}}{d\sigma_{\text{LO}}}$$

$$K_{\text{HO}} = \frac{d\sigma_{\text{NLOPS}} - d\sigma_{\text{NLO}}}{d\sigma_{\text{NLO}}}$$

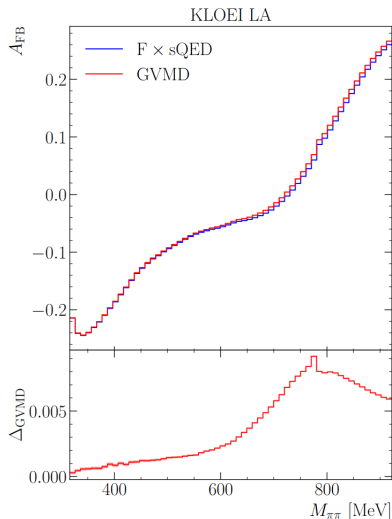


$M_{\pi\pi}, \theta_{\pi^+}$ - GVMD

$$K_{\text{GVMD}} = \frac{d\sigma_{\text{GVMD}} - d\sigma_{\text{F} \times \text{sQED}}}{d\sigma_{\text{F} \times \text{sQED}}}$$



KLOE-LA asymmetry



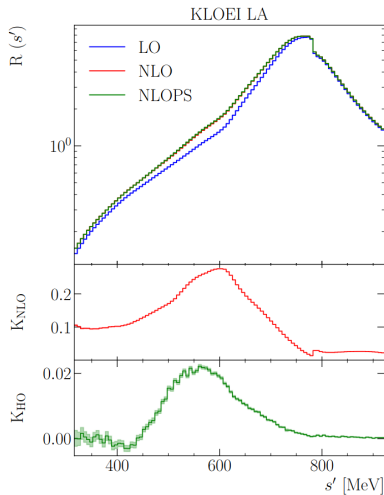
$$A_{FB}(M_{\pi\pi}) = \frac{d\sigma_F - d\sigma_B}{d\sigma_F + d\sigma_B},$$

where

$$d\sigma_F = \int_0^1 \frac{d\sigma}{dM_{\pi\pi} d\cos\theta^+} d\cos\theta^+,$$

$$d\sigma_B = \int_{-1}^0 \frac{d\sigma}{dM_{\pi\pi} d\cos\theta^+} d\cos\theta^+.$$

KLOE-LA R-ratio



$$R(s') = \frac{d\sigma(e^+e^- \rightarrow \pi^+\pi^-\gamma)/ds'}{d\sigma(e^+e^- \rightarrow \mu^+\mu^-\gamma)/ds'}$$

(a) KLOE I Large-Angle (LA)

$$\sqrt{s} = 1.02 \text{ GeV}$$

$$50^\circ \leq \theta^\pm \leq 130^\circ$$

$$p_z^\pm \geq 90 \text{ MeV} \vee p_\perp^\pm \geq 160 \text{ MeV}$$

$$50^\circ \leq \theta_\gamma \leq 130^\circ \wedge E_\gamma \geq 20 \text{ MeV}$$

$$0.1 \text{ GeV}^2 \leq M_{XX}^2 \leq 0.85 \text{ GeV}^2$$

(b) KLOE II Small-Angle (SA)

$$\sqrt{s} = 1.02 \text{ GeV}$$

$$50^\circ \leq \theta^\pm \leq 130^\circ$$

$$p_z^\pm \geq 90 \text{ MeV} \vee p_\perp^\pm \geq 160 \text{ MeV}$$

$$\theta_\gamma \leq 15^\circ \vee \theta_\gamma \geq 165^\circ$$

$$0.35 \text{ GeV}^2 \leq M_{XX}^2 \leq 0.95 \text{ GeV}^2$$

(c) BES III

$$\sqrt{s} = 4 \text{ GeV}$$

$$|\cos \theta^\pm| \leq 0.93 \wedge p_\perp^\pm \geq 300 \text{ MeV}$$

$$[|\cos \theta_\gamma| \leq 0.8 \wedge E_\gamma \geq 25 \text{ MeV}]$$

$$\vee [0.86 \leq |\cos \theta_\gamma| \leq 0.92 \wedge E_\gamma \geq 50 \text{ MeV}]$$

$$\exists! \gamma : E_\gamma \geq 400 \text{ MeV}$$

(d) B

$$\sqrt{s} = 10 \text{ GeV}$$

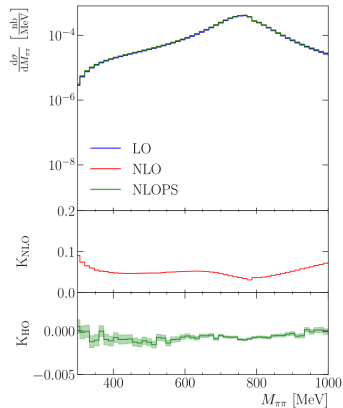
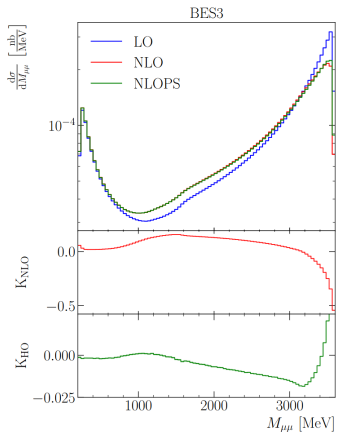
$$0.65 \text{ rad} \leq \theta^\pm \leq 2.75 \text{ rad} \wedge p^\pm \geq 1 \text{ GeV}$$

$$0.6 \text{ rad} \leq \theta_\gamma \leq 2.7 \text{ rad} \wedge E_\gamma \geq 3 \text{ GeV}$$

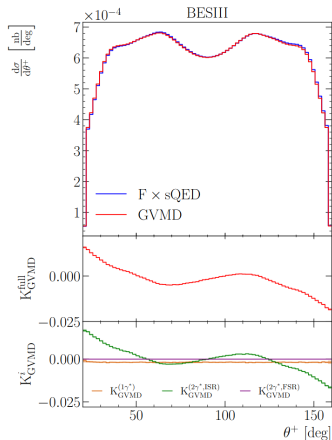
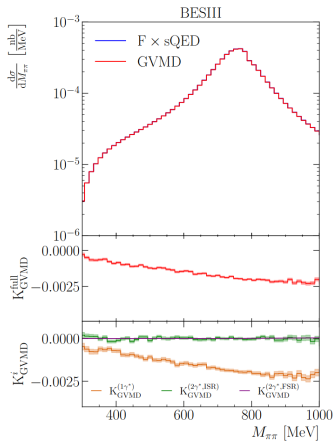
$$\theta_{\bar{\gamma}, \gamma^{(h)}} \leq 0.3 \text{ rad}$$

$$M_{XX\gamma} \geq 8 \text{ GeV}$$

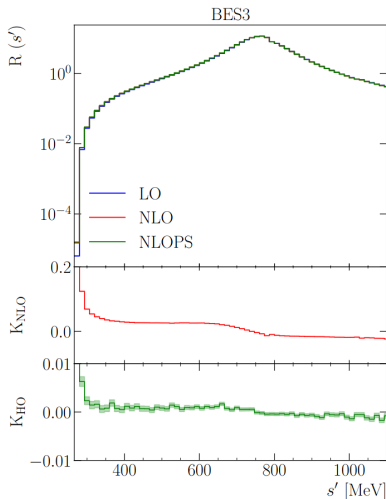
BESIII $M_{\pi\pi} - F \times s$ QED



BESIII $M_{\pi\pi}, \theta_{\pi^+}$ - GVMD



BESIII R-ratio

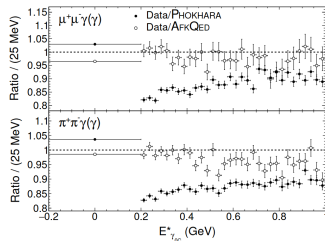


$$R(s') = \frac{d\sigma(e^+e^- \rightarrow \pi^+\pi^-\gamma)/ds'}{d\sigma(e^+e^- \rightarrow \mu^+\mu^-\gamma)/ds'}$$

BaBar test (arXiv:2308.05233)

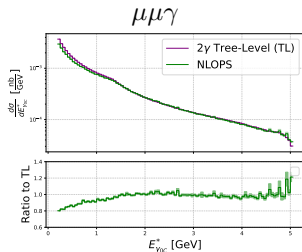
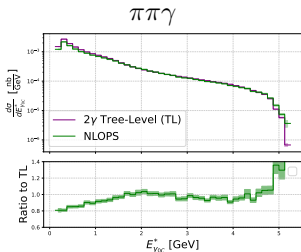
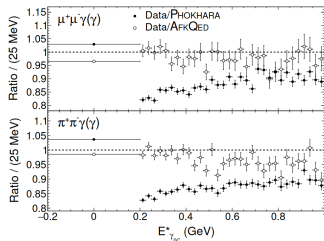


- $E_\gamma^* > 4 \text{ GeV}$ and $0.35 < \theta_\gamma < 2.4 \text{ rad}$
- $p_\perp > 0.1 \text{ GeV}$ and $0.4 < \theta_\pm < 2.45 \text{ rad}$
- $M_{\mu\mu} < 1.4 \text{ GeV}$ and $0.6 < M_{\pi\pi} < 0.9 \text{ GeV}$
- p_{0c} missing momentum with $E_{0c}^* > 200 \text{ MeV}$



BaBar test (arXiv:2308.05233)

- $E_\gamma^* > 4 \text{ GeV}$ and $0.35 < \theta_\gamma < 2.4 \text{ rad}$
- $p_\perp > 0.1 \text{ GeV}$ and $0.4 < \theta_\pm < 2.45 \text{ rad}$
- $M_{\mu\mu} < 1.4 \text{ GeV}$ and $0.6 < M_{\pi\pi} < 0.9 \text{ GeV}$
- p_{0c} missing momentum with $E_{0c}^* > 200 \text{ MeV}$



Summary and Outlook



- Multiple photon contributions are in general important for radiative return experiments.

Summary and Outlook



- Multiple photon contributions are in general important for radiative return experiments.
- GVMD structure dependent corrections can be of the same order of HO radiative corrections.
 - An independent study of a subset of GVMD corrections has appeared recently in [arxiv:2603.13171](https://arxiv.org/abs/2603.13171).

Summary and Outlook



- Multiple photon contributions are in general important for radiative return experiments.
- GVMD structure dependent corrections can be of the same order of HO radiative corrections.
 - An independent study of a subset of GVMD corrections has appeared recently in arxiv:2603.13171.
- This is part of a **common effort** (arxiv:2410.22882) to reduce and to better estimate theoretical uncertainties using precision MC generators: AfkQED, BabaYaga@NLO, KKMC, MCGPJ, McMule, Phokhara and Sherpa.

Summary and Outlook



- Multiple photon contributions are in general important for radiative return experiments.
- GVMD structure dependent corrections can be of the same order of HO radiative corrections.
 - An independent study of a subset of GVMD corrections has appeared recently in arxiv:2603.13171.
- This is part of a **common effort** (arxiv:2410.22882) to reduce and to better estimate theoretical uncertainties using precision MC generators: AfkQED, BabaYaga@NLO, KKMC, MCGPJ, McMule, Phokhara and Sherpa.
- **Future developments**
 - Complete $\pi\pi\gamma$ with FsQED.
 - $ee\gamma$ at NLOPS accuracy.
 - Extension to other hadronic channels ($K^+K^- \dots$)

Thank you !

