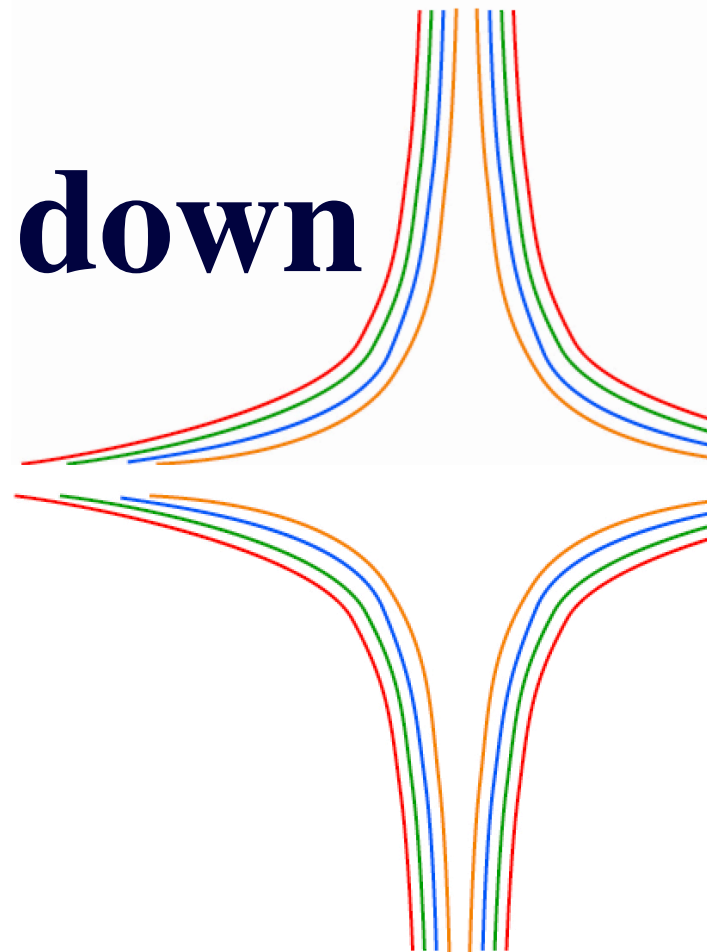


Spikes all the way down

Iosif Bena

IPhT, CEA

Université Paris-Saclay



Dimitrios Toulikas, Soumangsu Chakraborty, Raphaël Dulac, Zixia Wei,
Angèle Lochet, Yixuan Li, Nejc Čeplak, Shaun Hampton, Anthony Houppe,
Nick Warner, Pierre Heidmann, Antoine Bourget, Severin Lüst

erc

DE LA RECHERCHE À L'INDUSTRIE

cea

SACLAY

An amazing success of String Theory:

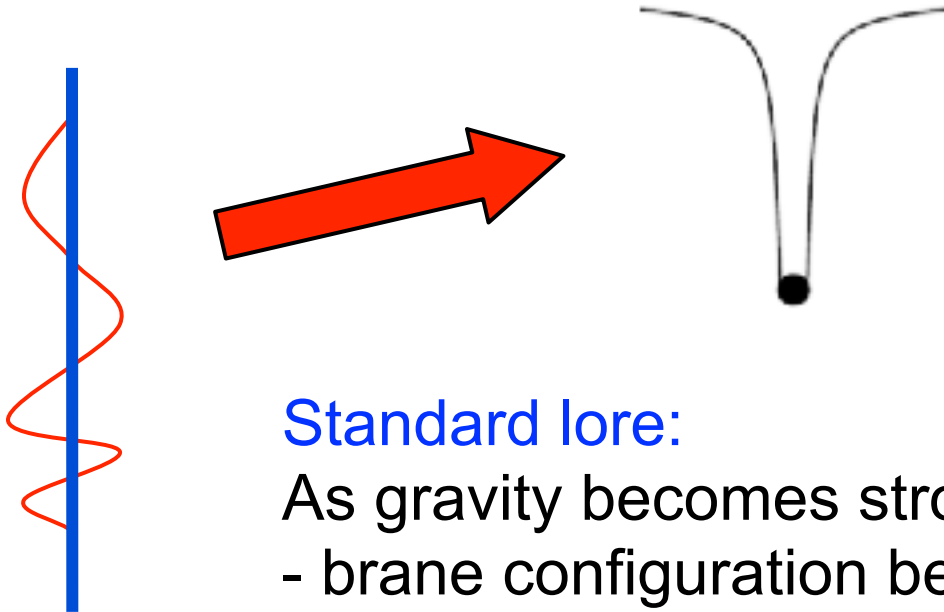
Count Black Hole Microstates (branes + strings)

Correctly match B.H. entropy !!!

Strominger, Vafa '96

Zero Gravity

One Particular Microstate at Finite Gravity:



Standard lore:

As gravity becomes stronger,

- brane configuration becomes smaller
- horizon develops and engulfs it
- recover standard black hole

Susskind
Horowitz, Polchinski
Chen, Maldacena, Witten

An amazing success of String Theory:

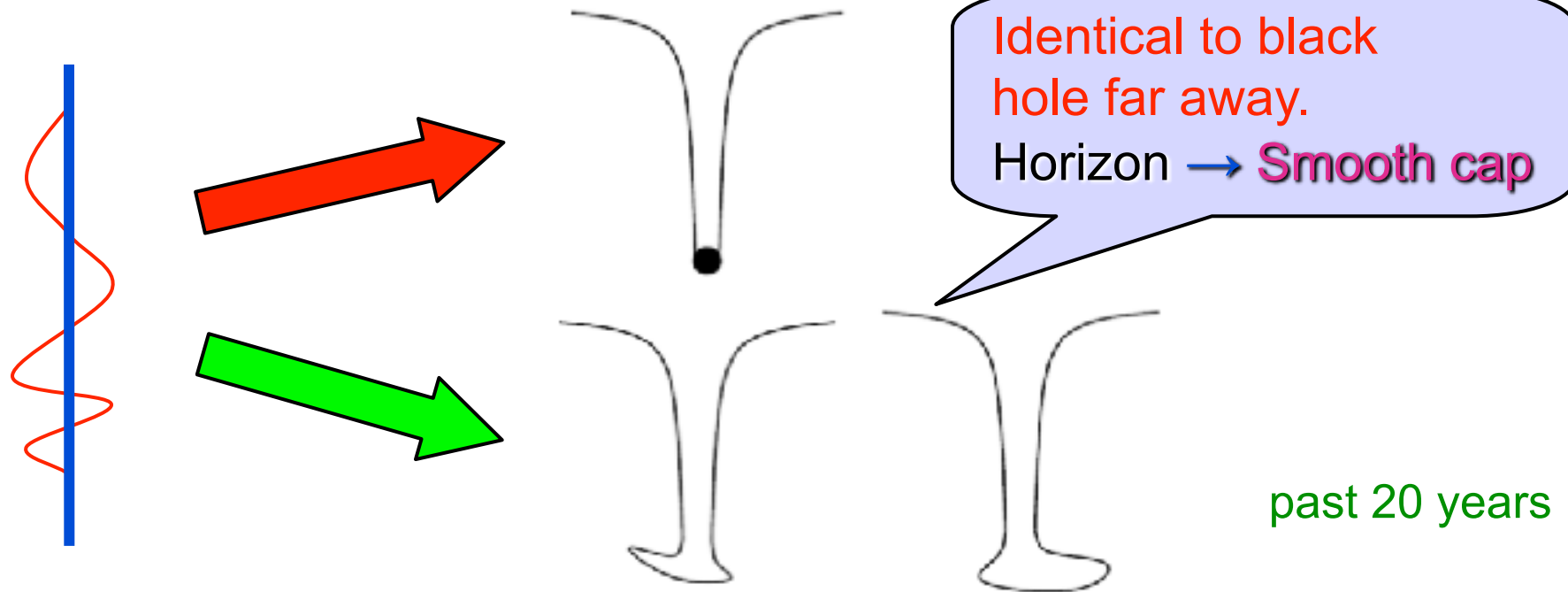
Strominger, Vafa '96

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One Particular Microstate at Finite Gravity:



BIG QUESTION: Are there enough geometries with no horizon to span BH Hilbert space ?

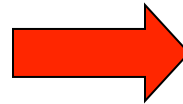
Analogy with ideal gas:

Thermodynamics

(Air = ideal fluid)

$$P V = n R T$$

$$dE = T dS + P dV$$



Statistical Physics

(Air -- molecules)

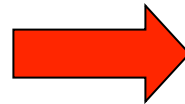
e^S microstates

typical

atypical

Thermodynamics

Black-Hole Solution



Statistical Physics

**Horizonless
microstate geometries**

Long distance physics
Gravitational lensing

Physics at horizon

Information loss
Gravity waves ?

Two routes

1. Build lots of solutions with black-hole mass and charges, but **no horizon**

2004-now

- **Bubbling geometries, superstrata**
- Many features of **typical** microstates
(mass gap = $2/N_1 N_5$)
- $S \sim (Q_1 Q_5)^{1/2} (Q_p)^{1/4} < S_{\text{BH}} \sim (Q_1 Q_5 Q_p)^{1/2}$ Mayerson, Shigemori '20
- Non-supersymmetric solutions
(essentially adding antibranes)

Heidmann, Bah '20-now

2. Track String-Theory microstates from no-gravity regime where they are counted

Best starting point: IIA F1-NS5

One F1 inside N_5 NS5 branes $\rightarrow N_5$ little strings.

Dijkgraaf, Verlinde, Verlinde

- Visible as **M2 brane strips** in M-theory
- **Total** $N_1 N_5$ independent **momentum carriers**
- each has **4 oscillation directions** (T^4) + **4 fermionic partners**

$$S = 2\pi\sqrt{\frac{4+2}{6}N_1N_5N_p} = S_{BH}$$

M2 along **y**, **11**

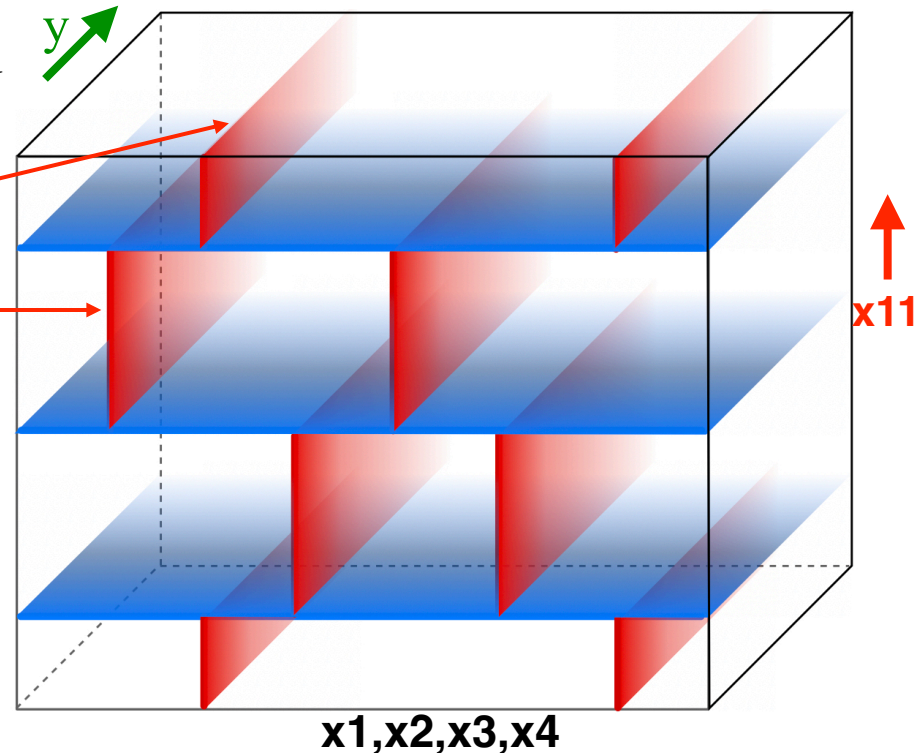
M5 along **y**, 1234

P along **y**

D1-D5: fractionated P

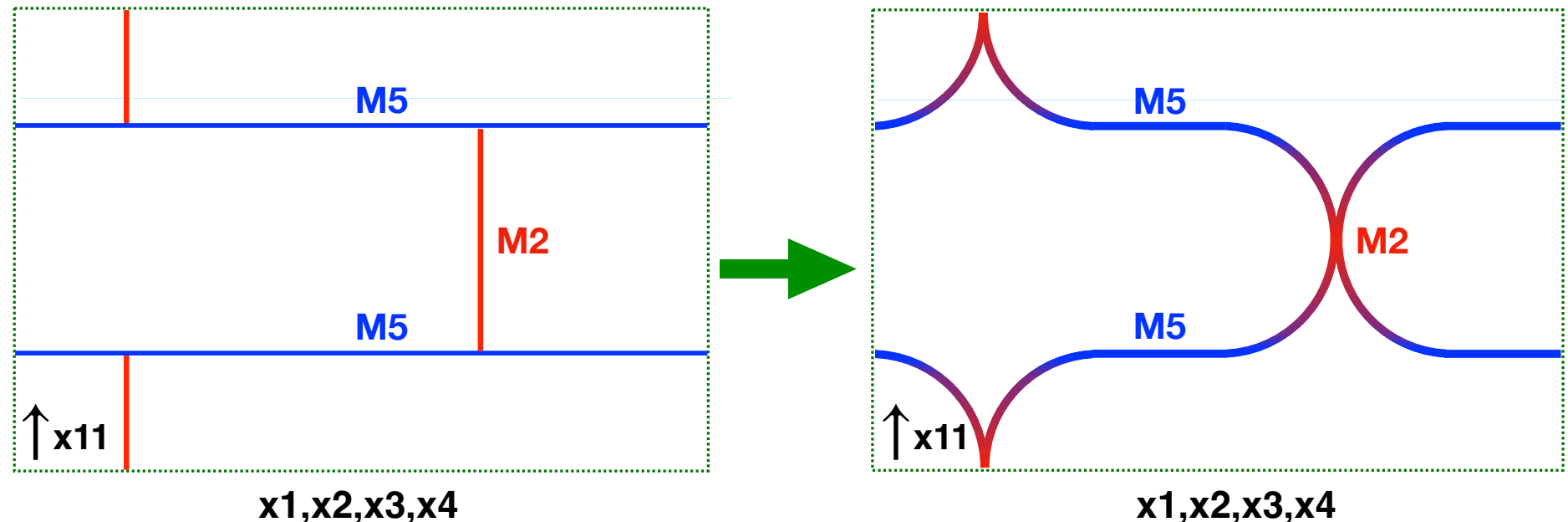
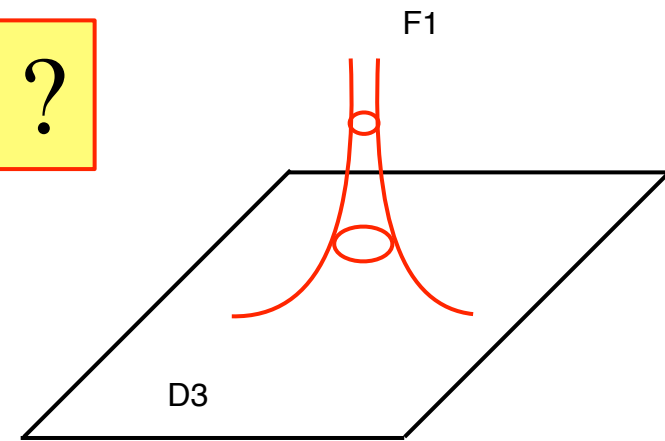
F1-NS5: fractionated F1

zero-coupling picture

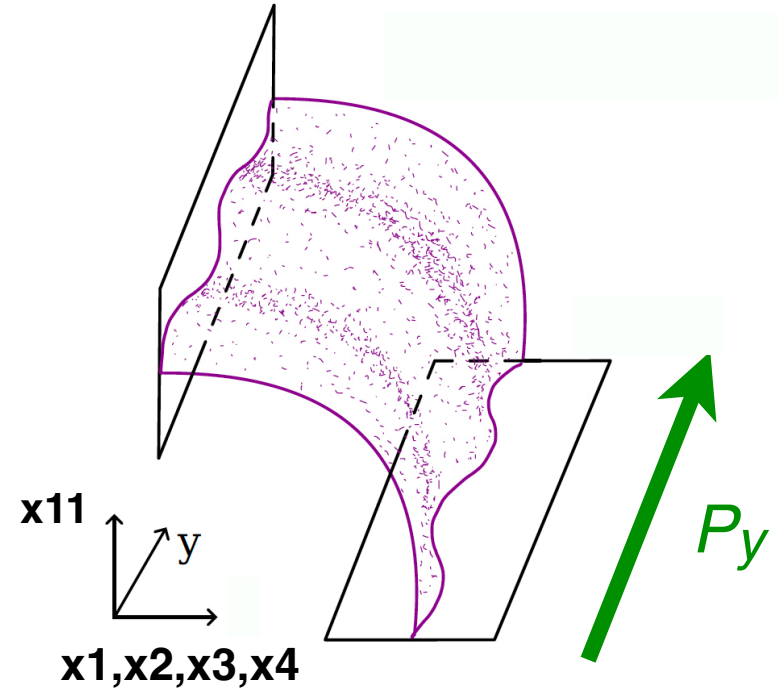
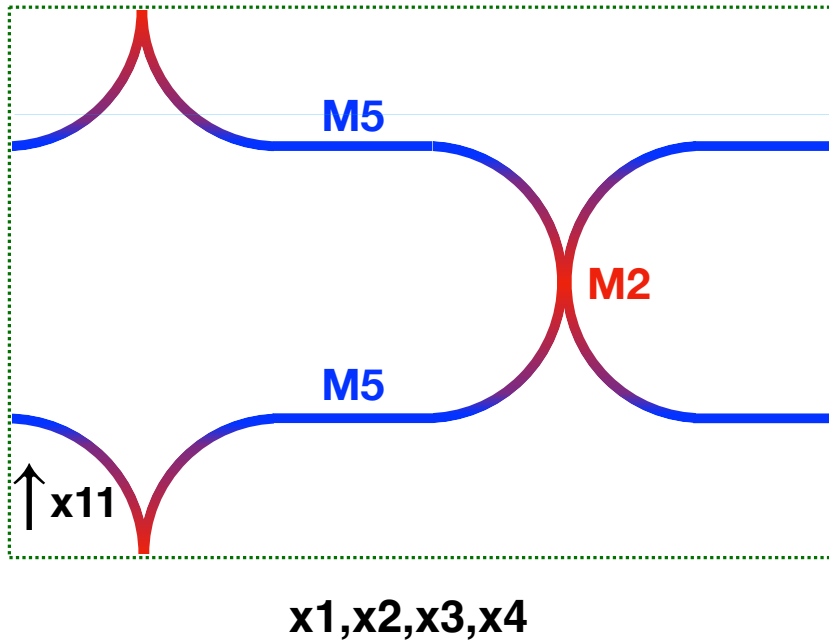


What about finite coupling ?

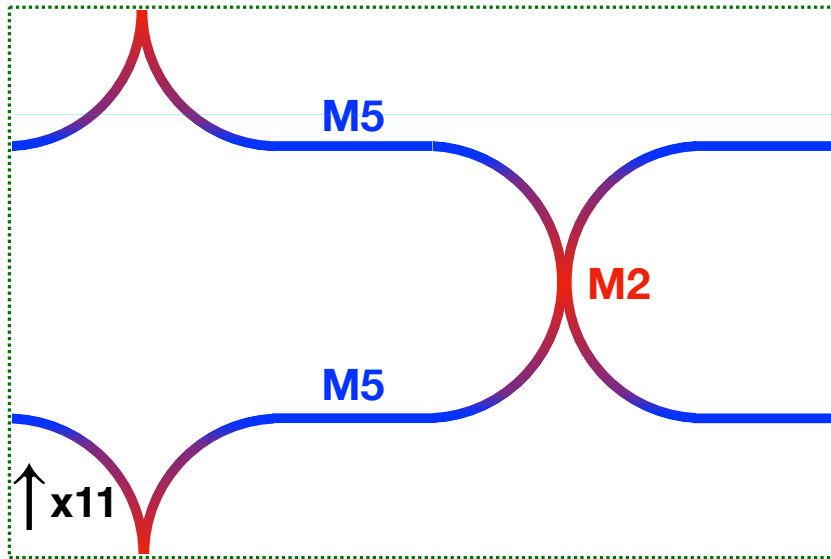
- Reminder:
Callan-Maldacena spike formed by F1 pulling on an orthogonal D3
- M2 branes also pull on the M5 brane
- Maze of supersymmetric branes: [super-maze](#)



Spike \rightarrow *furrow* carrying momentum waves along common M2-M5 direction (y)



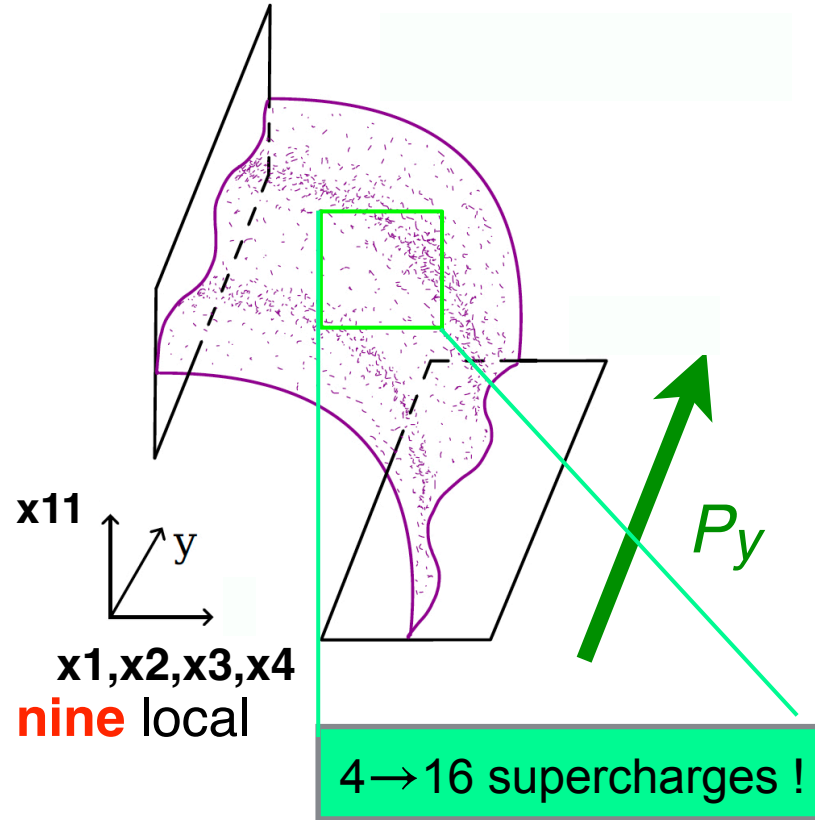
Spike \rightarrow *furrow* carrying momentum waves
along common M2-M5 direction (y)



x_1, x_2, x_3, x_4

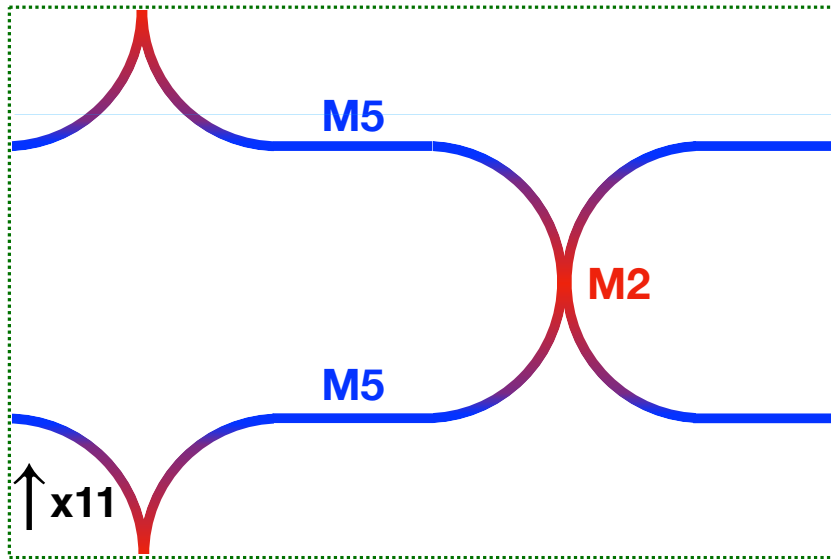
Zoom in on the furrow carrying momentum: **nine** local
brane charges: $M2_{x_{11},y}$ $M5_{y,x_1,x_2,x_3,x_4}$ P_y

$M2_{x_1,x_{11}}$ $M5_{x_{11},y,x_2,x_3,x_4}$ $M2_{x_1,y}$ $M5_{x_{11},x_1,x_2,x_3,x_4}$ $P_{x_{11}}$ P_{x_1}



Bena, Hampton, Houppe, Li, Toulikas

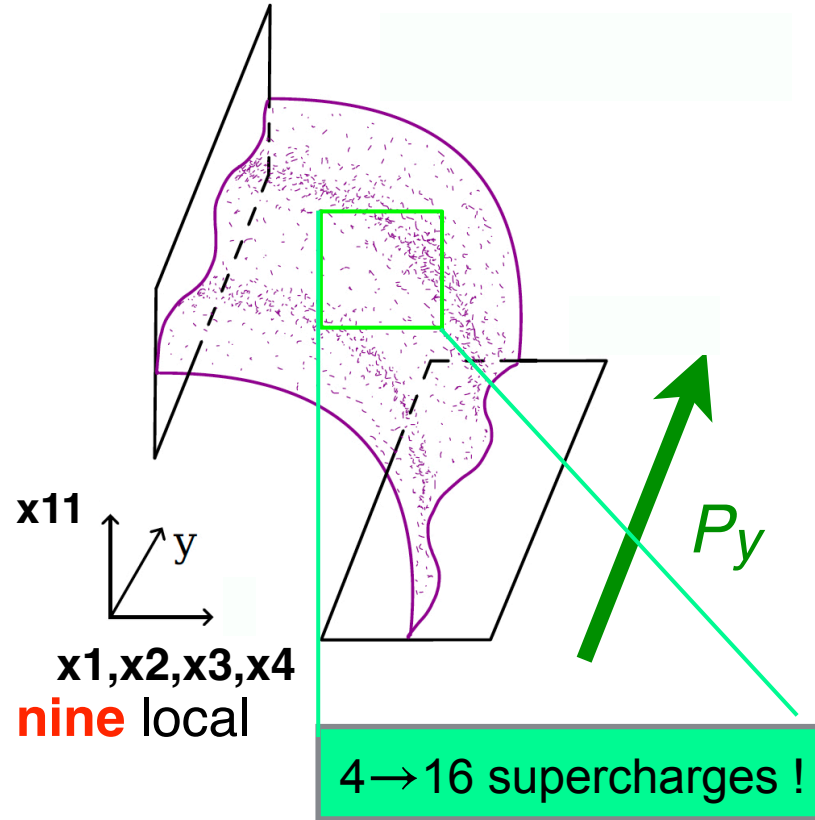
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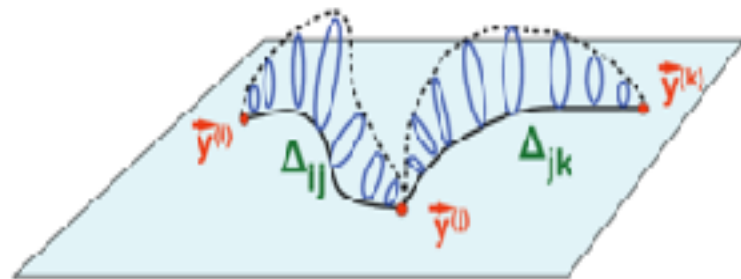


Bena, Hampton, Houppe, Li, Toulukas

Smoking gun of smooth horizonless solutions

A bit of history

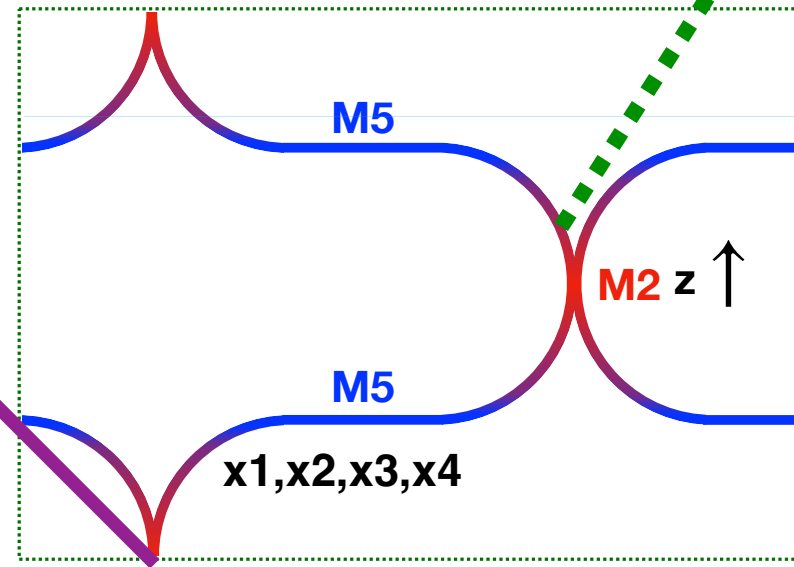
- First microstate geometries
 - Horizonless bubbling solutions Bena, Warner '06
 - \Leftrightarrow Multicenter fluxed D6 branes Balasubramanian & al '06
 - **16 susy** at every center, **4 globally**
 - Entropy much smaller than BH de Boer & friends
- Microstate geometries with supertubes
 - Functions of one variable Bena, Bobev, Giusto, Ruef, Warner '10
 - Smooth \Leftrightarrow **16 susy** when zooming on supertube
- Superstrata. conjectured in Bena, de Boer, Shigemori, Warner '11
 - Fns. of 2 variables; **16 susy locally**, **4 globally**
 - **HABEMUS**: Smooth. Bena, Giusto, Russo, Shigemori, Warner '15
- Smooth horizonless sols \Leftrightarrow brane config. with **16 susy locally**, **4 globally**
- **Big Goal: build the super-maze solutions**



Add another type of brane

- + $M5'$ - still 8 susy
- bad for BH microstates (space filling)
- Generic solution describes:
 - $M2$ suspended between $M5$ and $M5'$
 - infinite $M2$ spikes ending on $M5$ or on $M5'$
 - $M2$ between two $M5$ or two $M5'$
 - $M2$ crossing $M5$ or $M5'$ but not ending on them

$M5'$ x_5, x_6, x_7, x_8



Monge-Ampère equation:

$$\mathcal{L}_v G_0 = (\mathcal{L}_u G_0) (\partial_z \partial_z G_0) - (\nabla_{\vec{u}} \partial_z G_0) \cdot (\nabla_{\vec{u}} \partial_z G_0)$$

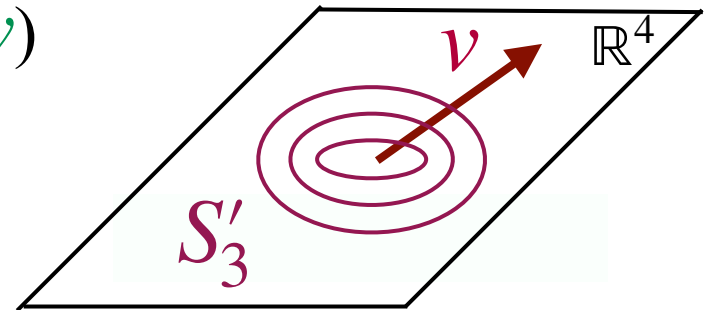
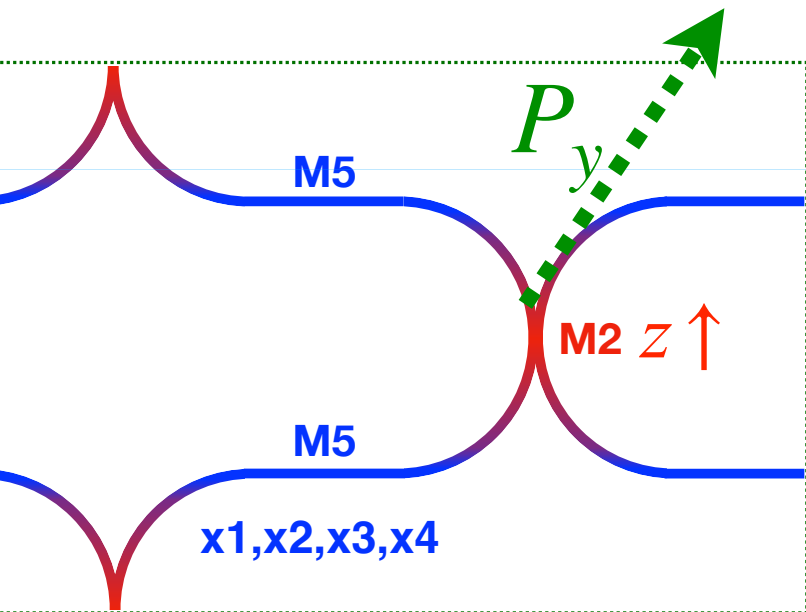
Simplest relevant solution: $SO(4) \times SO(4) \times U(1)_y$
 invariant on $S^3 \in M5$, on $S^3 \in M5'$, no P_y

- **Monge-Ampère:** $G_0(z, u, v)$ - hard
- Brane sources $\Rightarrow \exists$ solution (singular) Lunin
- **Black-hole charges** $M2_{zy} + M5_{1234y} + P_y \Rightarrow$
 at least *cohomogeneity-4* (z, u, v, y)

M2: 0, y , z

M5: 0, y , u , S_3

M5': 0, y , v , S'_3



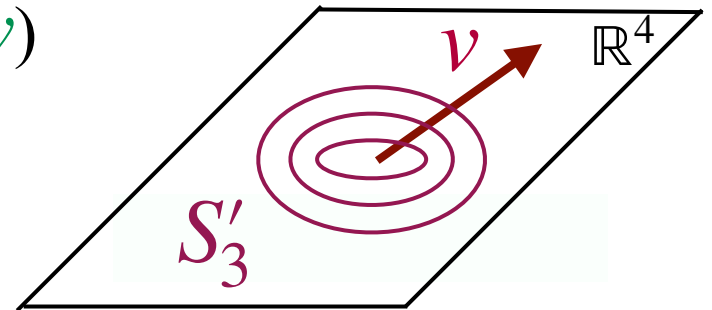
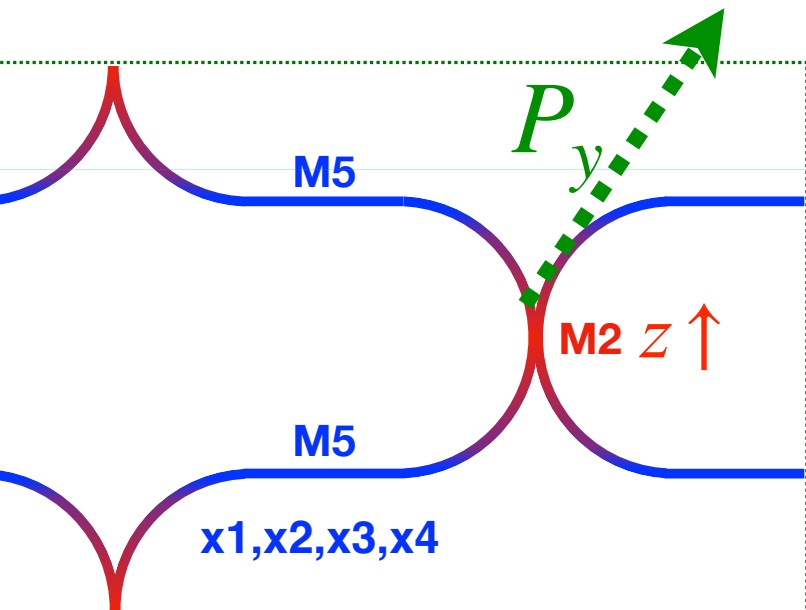
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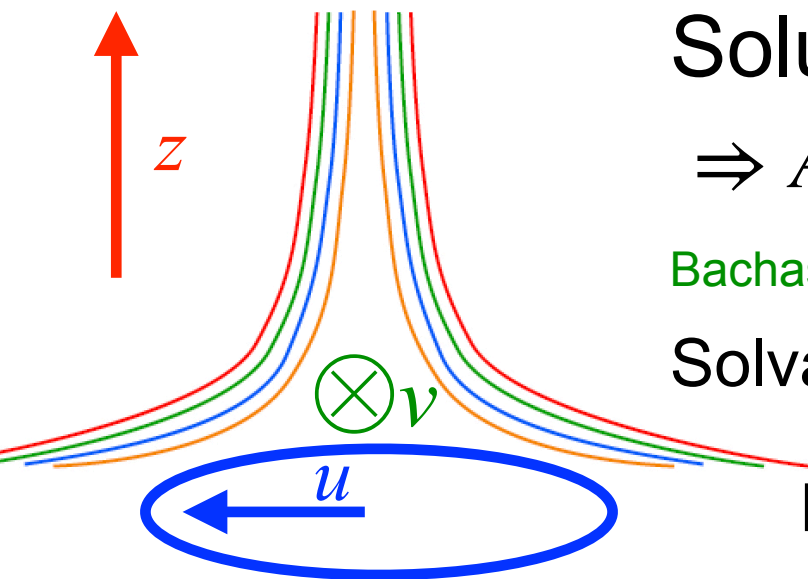
M5': 0, y, v, S'_3



But ... To the blind bird God
 sometimes makes the nest

Romanian proverb

Zoom in on brane profile



Solution with $SO(4) \times SO(4) \times \mathbb{R}^{1,1}$
 $\Rightarrow AdS_3 \times S^3 \times S^3 \times$ Riemann surface

Bachas, Estes, D'Hoker, Krym Cohom $3 \rightarrow 2$!

Solvable in a **linear** algorithm !

Brane coordinates: M2: 0, y, z

M5: 0, y, u , S_3

M5': 0, y, v , S'_3

$$ds_{AdS_3}^2 = \mu^2(-dt^2 + dy^2) + \frac{d\mu^2}{\mu^2}$$

μ + Riemann surface coordinates: x, y = functions of z, u, v

$AdS_3 \times S^3 \times S^3$ solution = backreaction of M2-M5 spikes

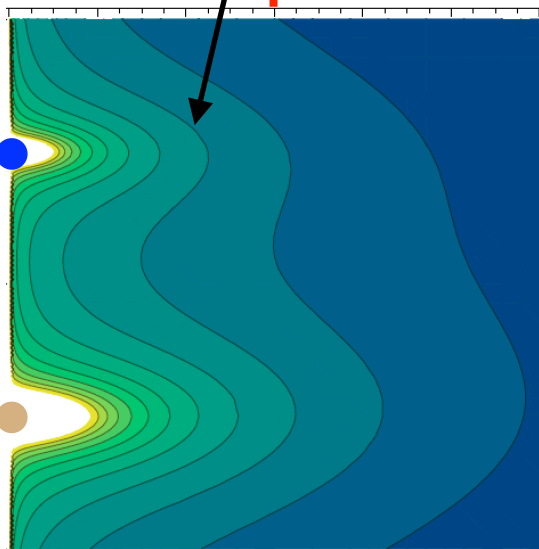
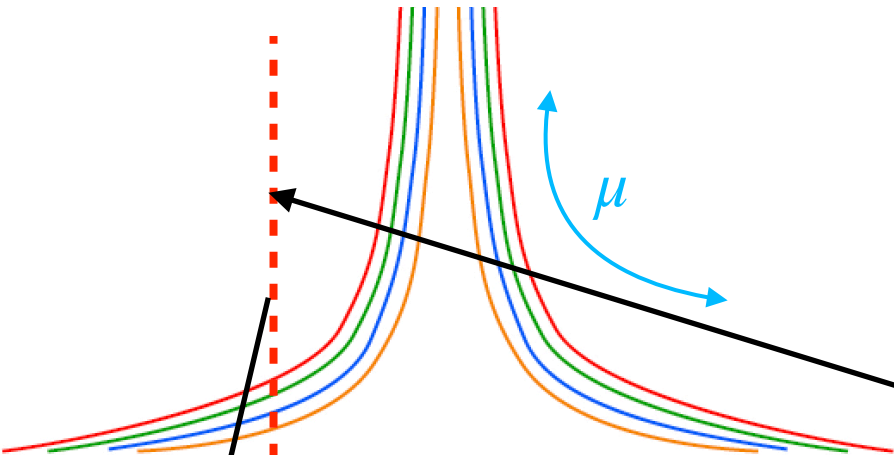
Bena, Chakraborty, Houppe, Toulukas, Warner

Zoom in on brane profile

Point on Riemann surface:

$z u^2 = \text{constant}$. M2-M5 spike
spanned by μ

M2 along z , at constant u , $v =$
curve on Riemann surface
 $x(\mu)$, $y(\mu)$



Zoom in on brane profile

Point on Riemann surface:

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M2 along z , at constant u , $v =$
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 $x(\mu)$, $y(\mu)$

Born-Infeld construction of superm2 +
null momentum wave Bena, Dulac

Momentum wave can be added to the
sugra solution following linear algorithm !
Bena, Dulac, Houppé, Toulakas, Warner

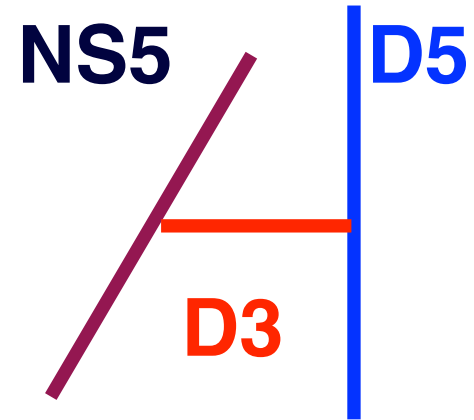
+ Arbitrary function of null coordinate \Rightarrow cohomogeneity 2 !!!

A similar system

D3: 0,1,2, z

D5: 0,1,2, u , S_2

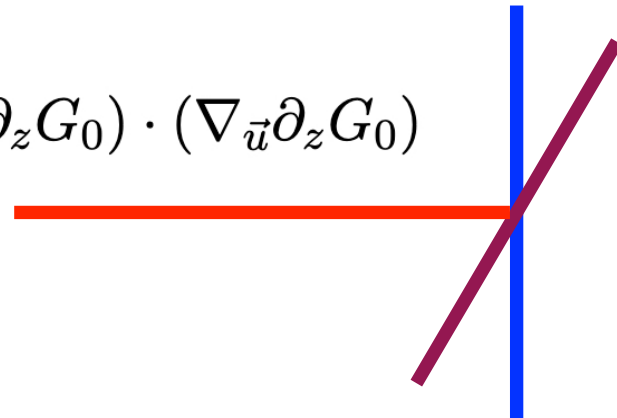
NS5: 0,1,2, v , S'_2



- Generic solution describes:
 - **D3** suspended between **D5** and **NS5**
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 - **D3** between two **D5** or two **NS5**
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Link with AdS_4 solutions

D3-D5-NS5 solution with $SO(3) \times SO(3) \times \mathbb{R}^{2,1}$
 (Monge-Ampère) $\Rightarrow AdS_4 \times S^2 \times S^2 \times$ Riemann surface

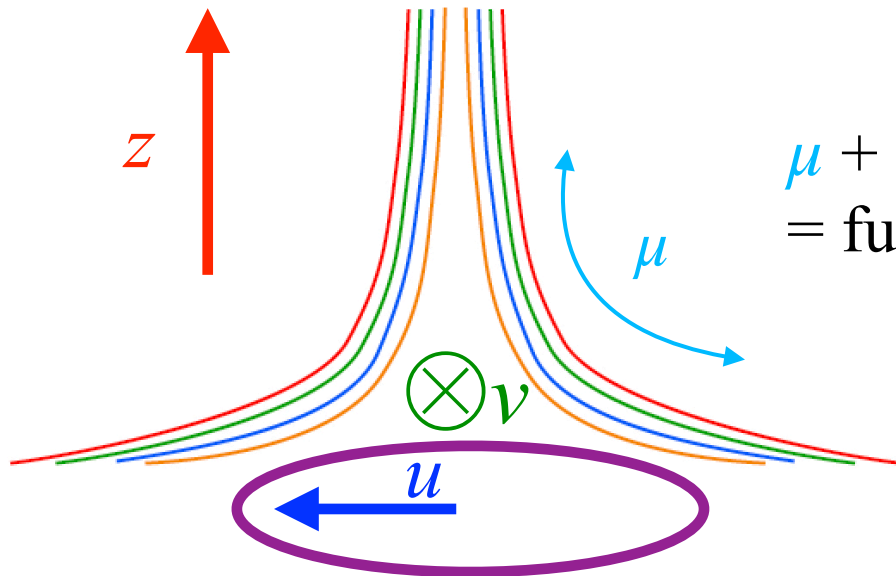
Brane coordinates: **D3**: 012, z

D5: 012, u , S_2

NS5: 012, v , S'_2

$$ds_{AdS_4}^2 = \mu^2(-dt^2 + dx_1^2 + dx_2^2) + \frac{d\mu^2}{\mu^2}$$

μ + Riemann-surface coordinates x, y
 = functions of z, u, v



Point on Riemann surface:

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 spanned by μ

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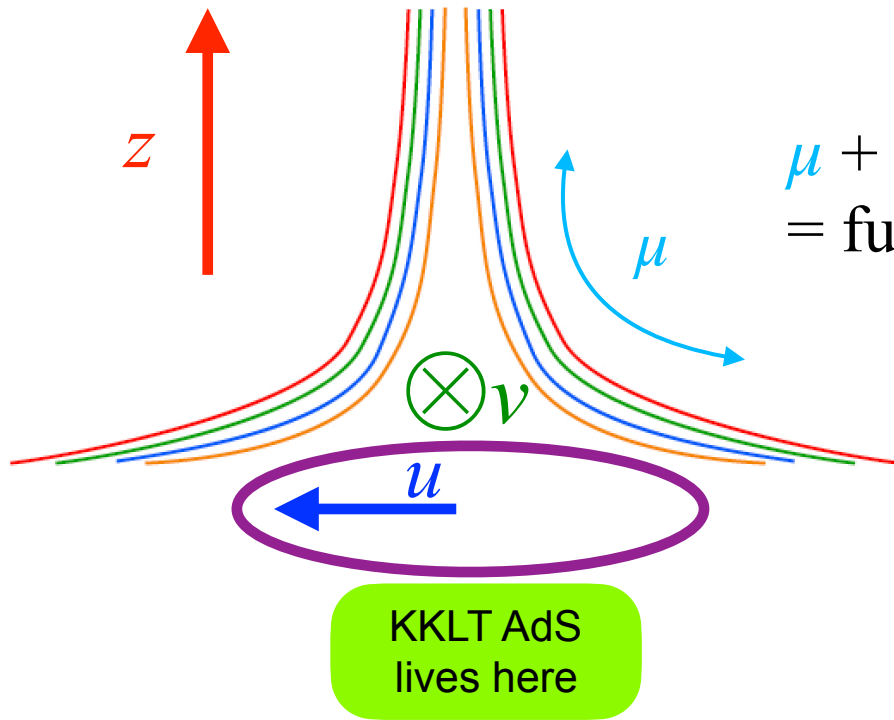
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KKLT = hot-potato field (*sulfureux*)

Compactify to 4D on 6D manifold (Calabi-Yau)

Lots of unphysical massless scalars (moduli)

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Compactify to **4D** on **6D manifold** (Calabi-Yau)

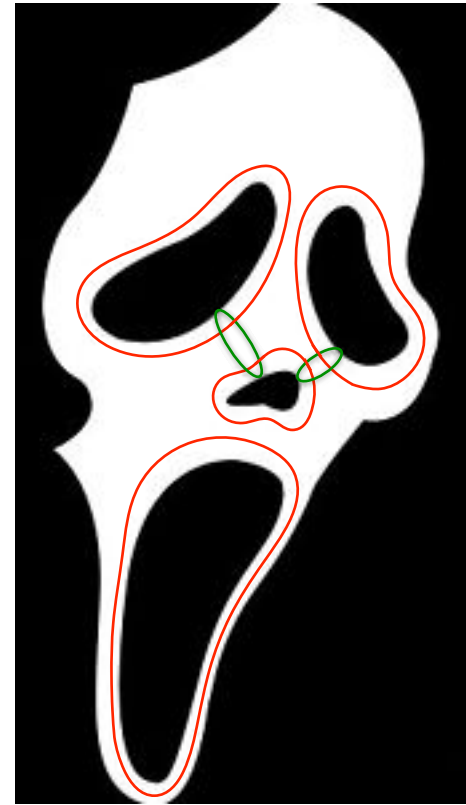
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1. **add fluxes** \rightarrow fix complex-str. moduli
2. gaugino cond. \rightarrow fix Kähler moduli
 \rightarrow **10^{500}** stable **AdS** vacua

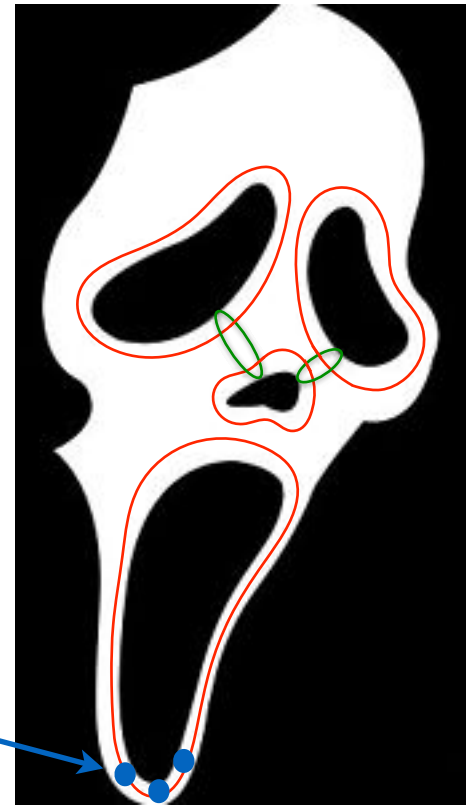


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3. **anti-D3** down **long** throats →
redshift → very-small **energy** →
lift **AdS** to **de Sitter** KKLT + 3500 others



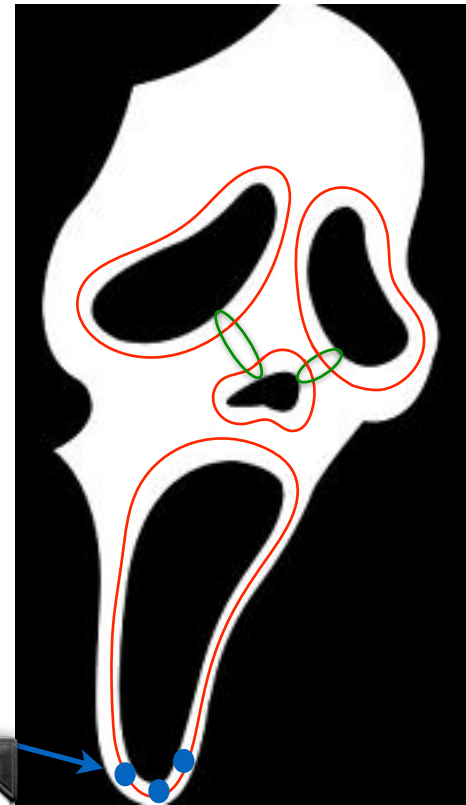
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THE LANDSCAPE



Why the sulfurocity ?

Steps 1,2,3: **low-energy effective field theory**
using String-Theory-derived **ingredients**

Nontrivial interactions in String Theory



Runaways



1

Bena, Dudaş, Graña, S. Lüst



Tadpole problem



1

Bena, Blåbäck, Graña, S. Lüst
Bena, Brodie, Graña

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+ lots of swampland-type conjectures:

scale separated dS and AdS ***not possible***

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hard to convince KKLT fans ...

Do not pray to the saint who does not help you !

Romanian proverb

The KKLT domain wall

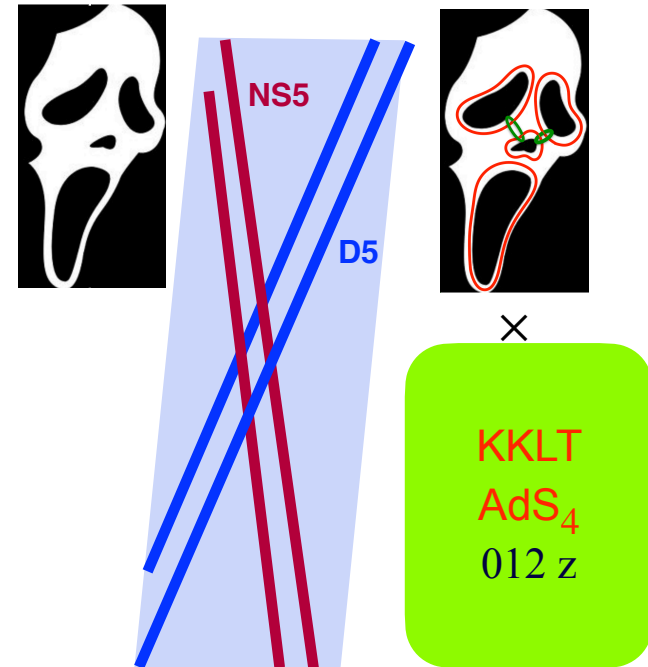
S.Lüst, Vafa, Wiesner, Xu

- Compactification fluxes eaten by branes on dual cycles

- $F_{\text{UV domain wall}} > \left(\frac{\ell_{\text{AdS}_4}^{\text{KKLT}}}{l_{\text{Planck}}^{4D}} \right)^2$

- Bound on $\ell_{\text{AdS}_4}^{\text{KKLT}}$

$$F_{\text{UV}}^{\text{perturbative}} \sim N^2$$



D5: $012 \times \Omega_3$

NS5: $012 \times \Omega'_3$

Domain wall: $012 \times \text{CY}$

The KKLT domain wall

S.Lüst, Vafa, Wiesner, Xu

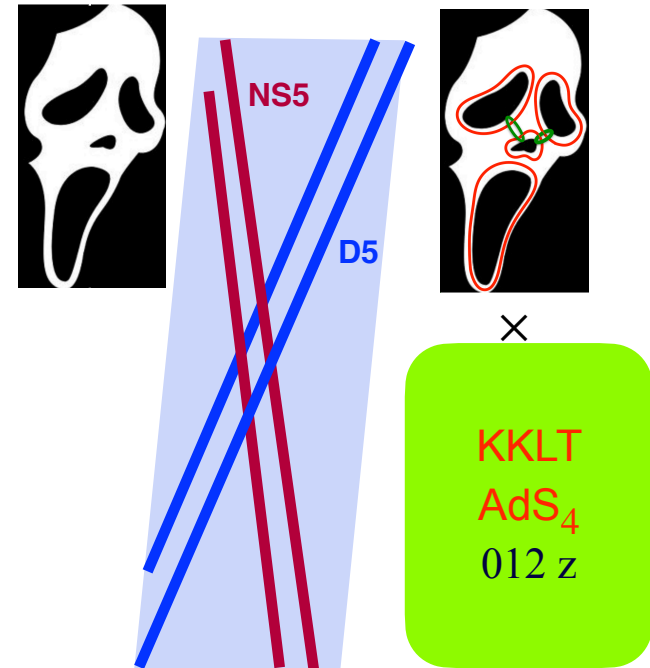
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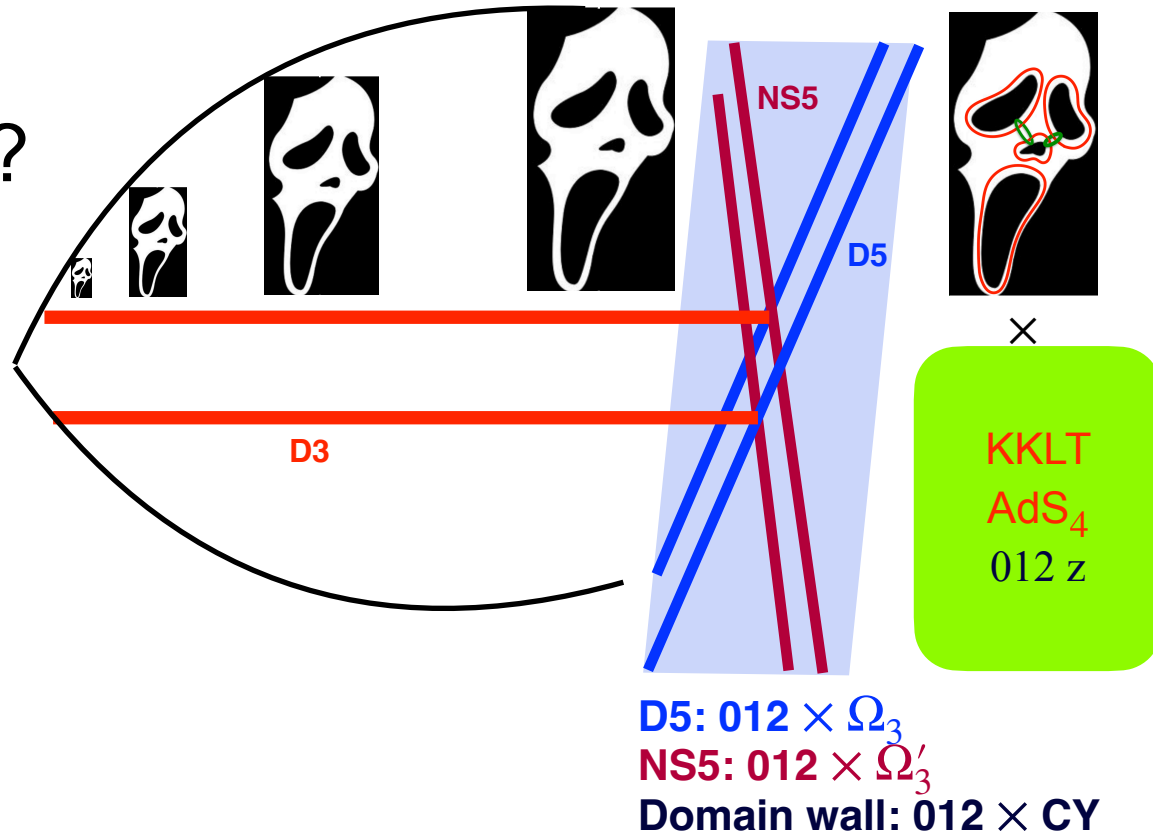
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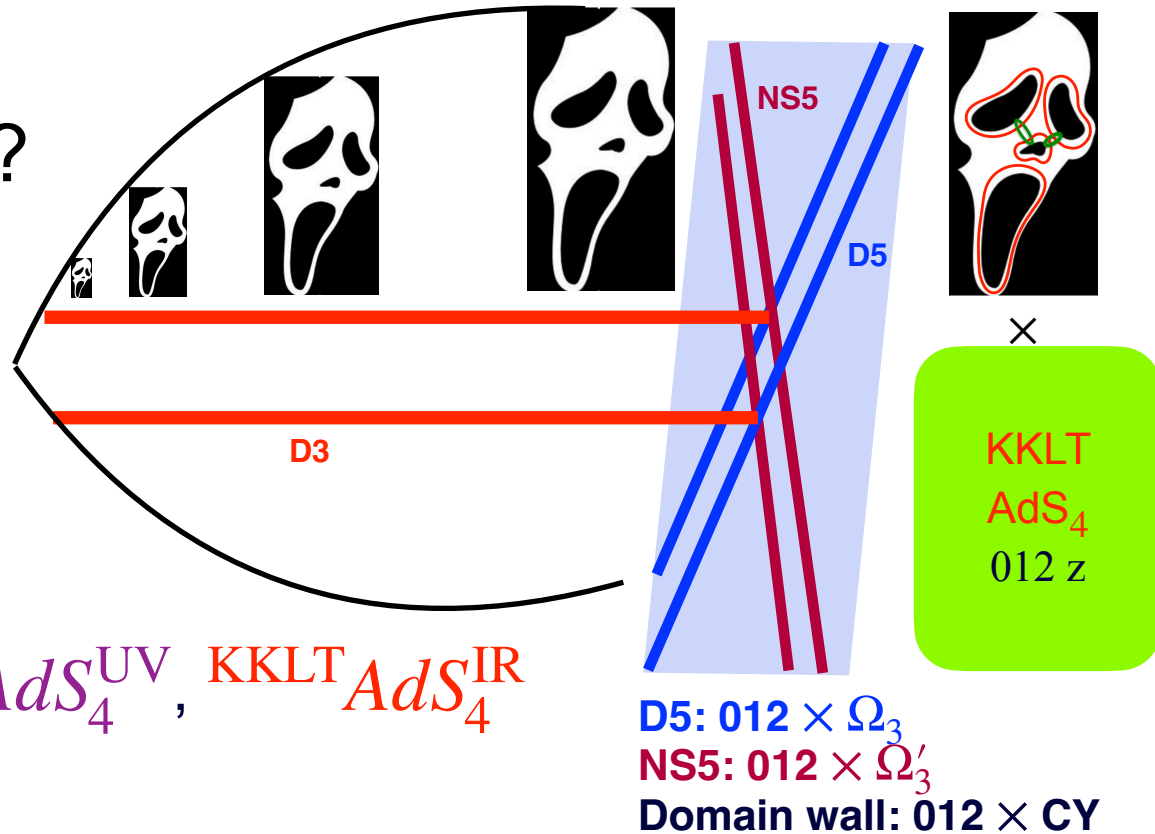
What is on the left ?

- CY shrinks
- Universe ends
- *KKLT ex nihilo*
Bena, Li, S.Lüst
- D3-D5-NS5 spikes

- Spike backreaction $\Rightarrow AdS_4^{\text{UV}}, \text{KKLT } AdS_4^{\text{IR}}$

- $\ell_{\text{AdS}_4}^{\text{UV}} > \ell_{\text{AdS}_4}^{\text{KKLT}}$

- $F_{\text{CFT}}^{\text{UV}} \sim N^4 \ni$ all domain wall d.o.f. Not enough for KKLT AdS_4

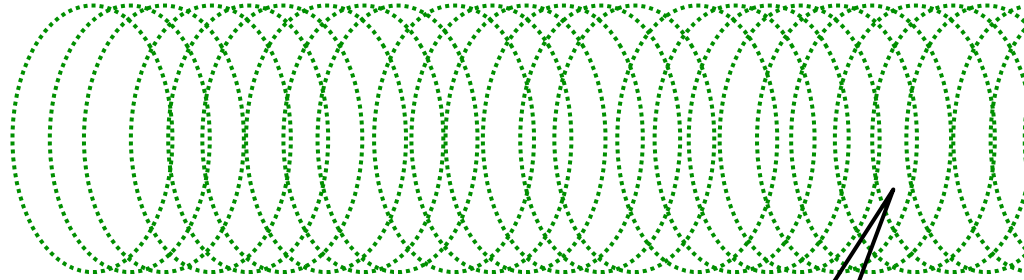


$AdS_4 \times S^2 \times S^2 \times \Sigma$ solutions: 3 classes

1. Two AdS_5 ends: Janus

NS5 smeared on $S^2 \times z^+$
codimension-1 for $z > 0$

$$\Phi = \Phi_1$$



deep inside $z \gg 0$:
 $\Phi = \Phi_2 = \text{constant}$
 $H_3 = 0$

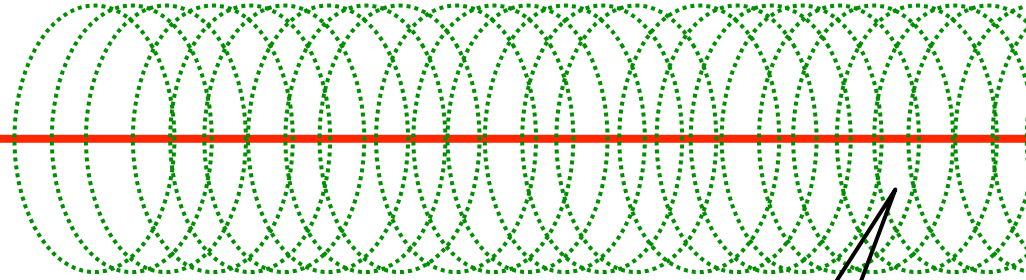
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D3

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$AdS_4 \times S^2 \times S^2 \times \Sigma$ solutions: 3 classes

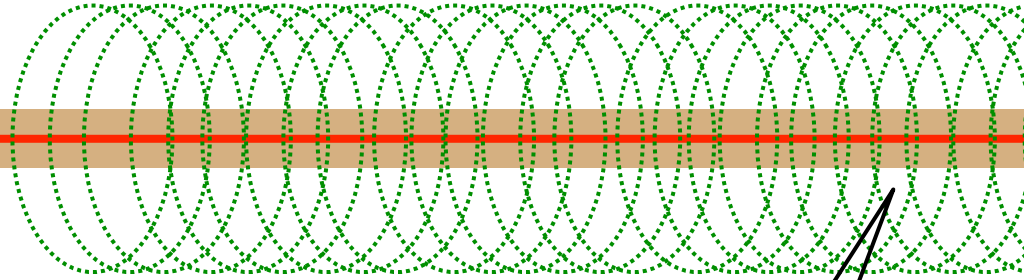
1. Two AdS_5 ends: **Janus**

NS5 smeared on $S^2 \times z^+$
codimension-1 for $z > 0$

AdS'_5

D3

AdS_5 $\Phi = \Phi_1$



deep inside $z \gg 0$:
 $\Phi = \Phi_2 = \text{constant}$
 $H_3 = 0$

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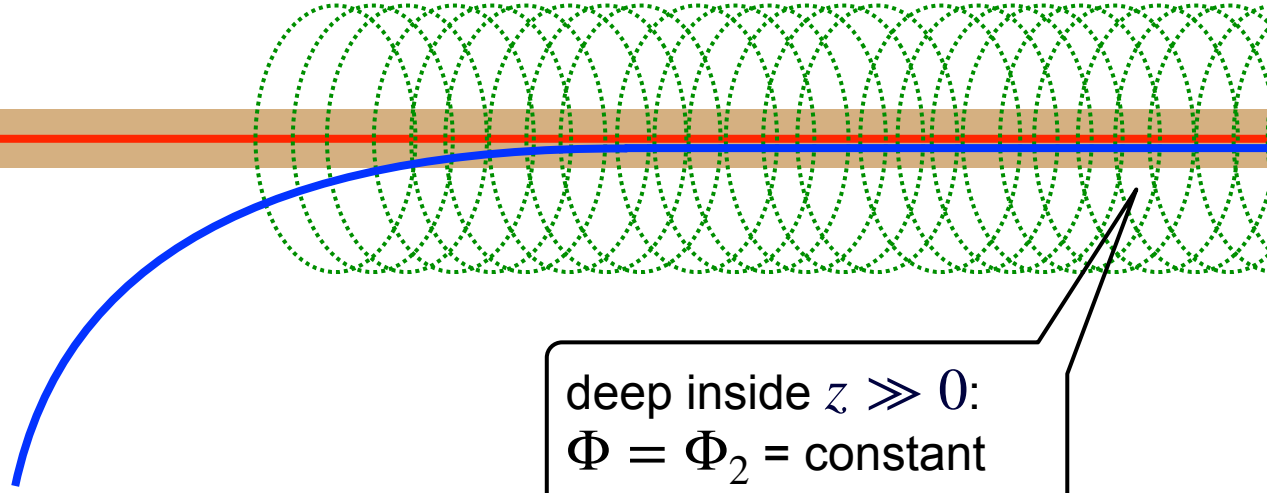
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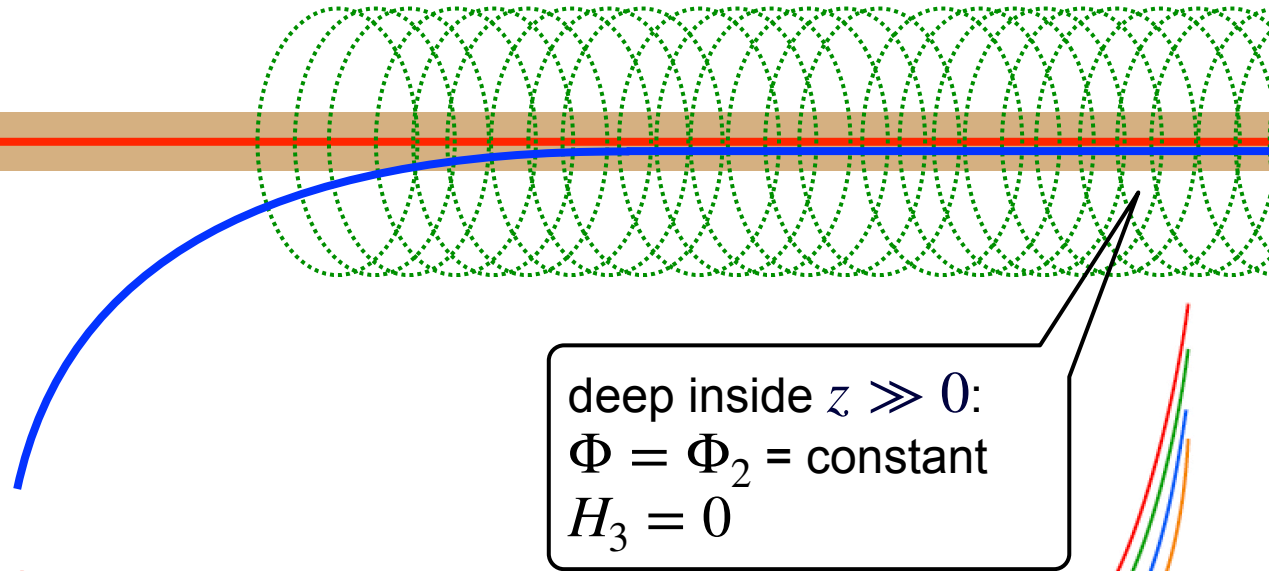
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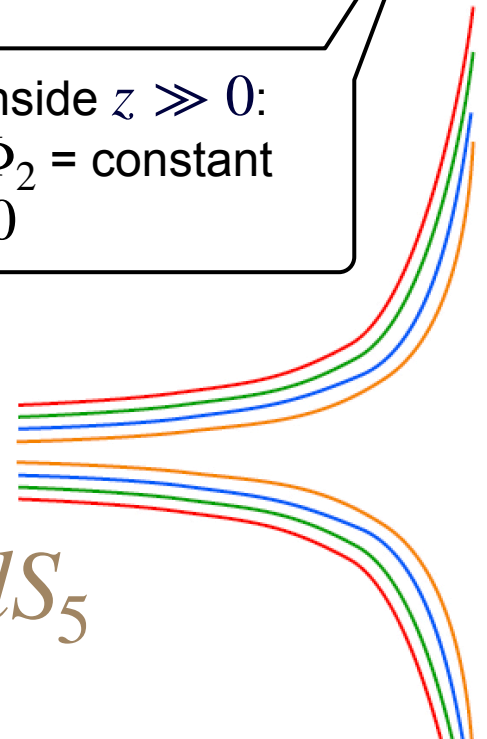


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2. One AdS_5 end:

D3 spikes on D5 and NS5 branes:

AdS_5



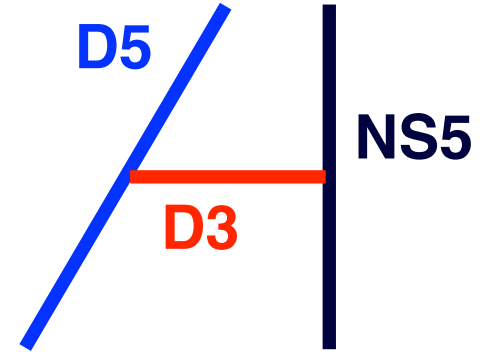
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Everybody and their brother is confused

Claim:

Directly ruled out by our mapping

μ direction parameterizes an **infinite spike**



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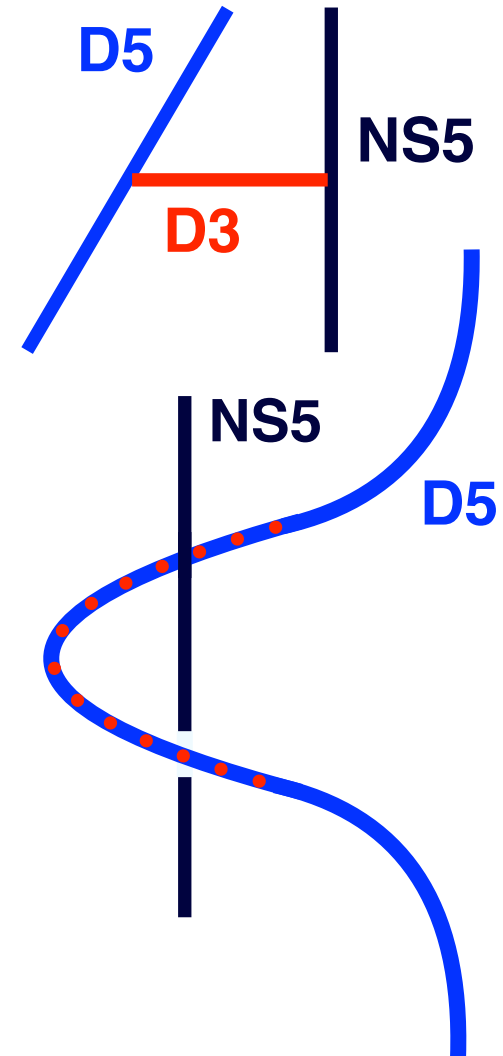
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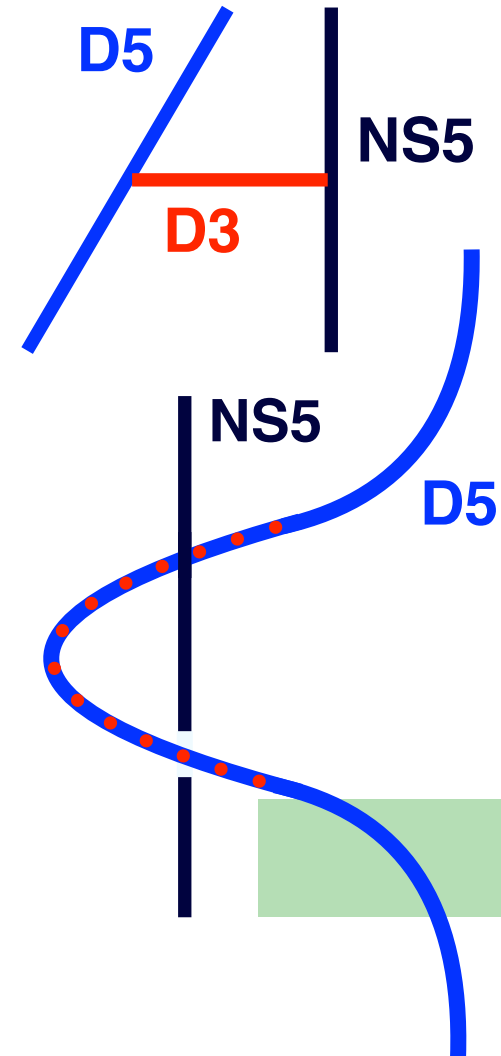
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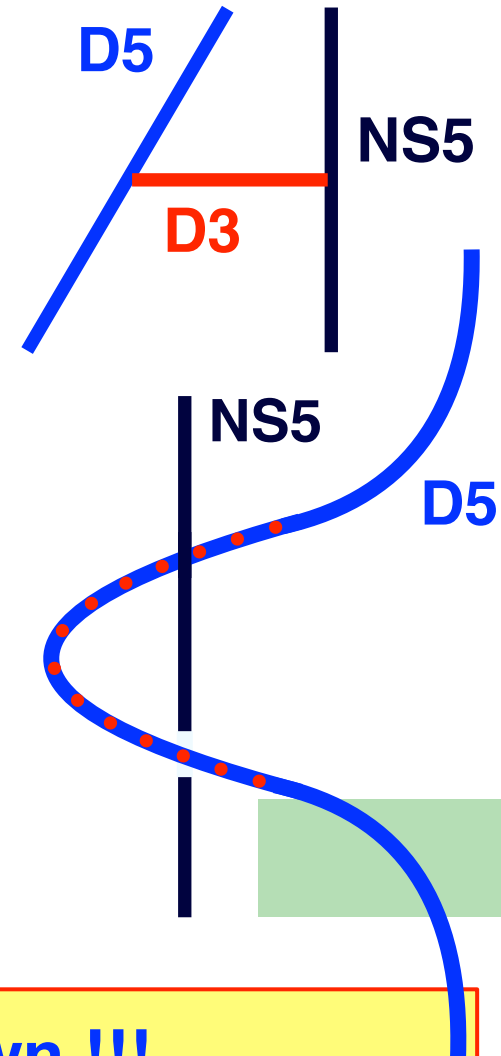
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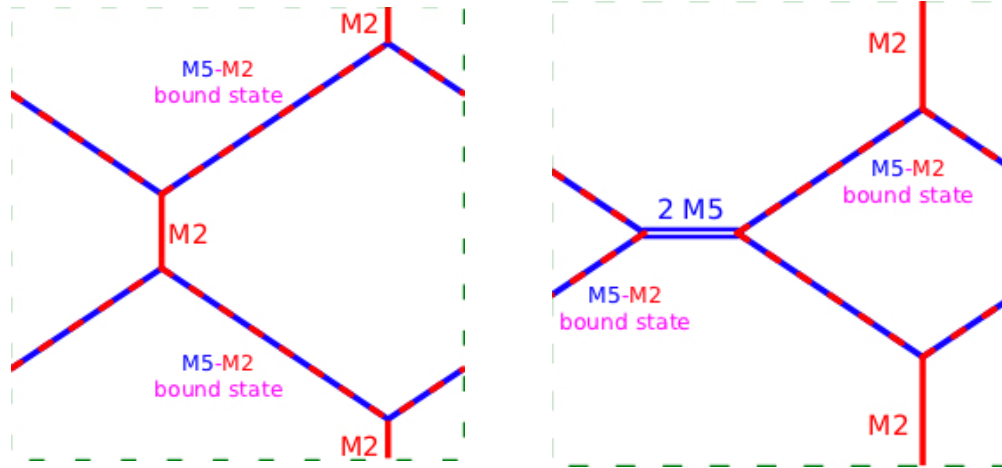
looks like a spike



AdS solutions : It's spikes all the way down !!!

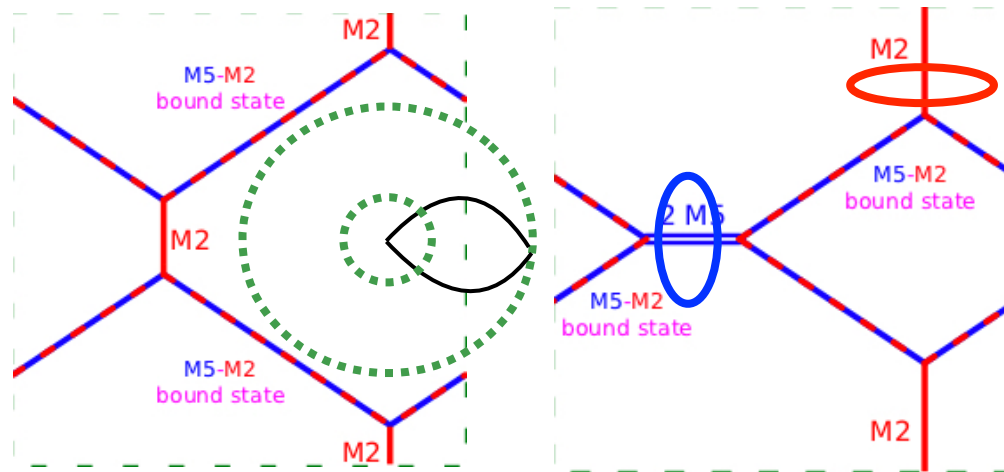
No chance to describe **D3** sandwiched between 2 **D5**'s ('t Hooft Polyakov monopole) or **D3** between **D5** and NS5
God makes to the nest of blind bird ... but very spiky nest

How will the $SO(4)$ -invariant solution look like ?



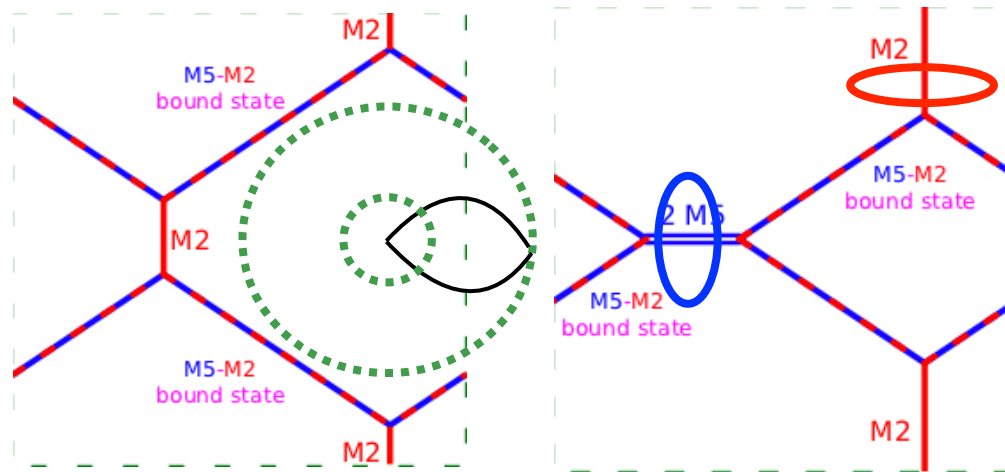
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- *16-susy locally \Rightarrow no horizon*
- Branes wrapping compact contractible cycles \Rightarrow **Geometric transition** \Rightarrow Bubbles wrapped by fluxes inside the internal dimensions.
- Smooth bubbling sources: currently trying to construct
- Expectation based on earlier work:
 - **backreaction** will make bubbles **large**
 - *irrespective* of T^4 size at infinity

The big hope: Track **each and every** BH microstate from zero-gravity regime to fully-backreacted solution

DVV microstates

$$S = S_{\text{BH}}$$

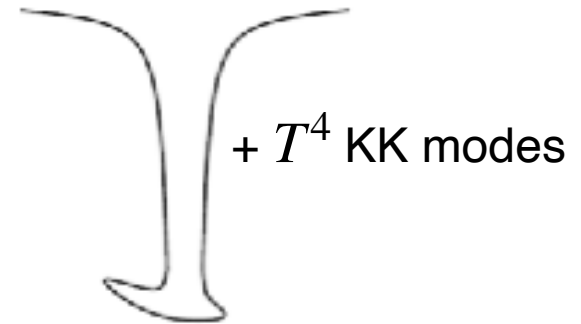
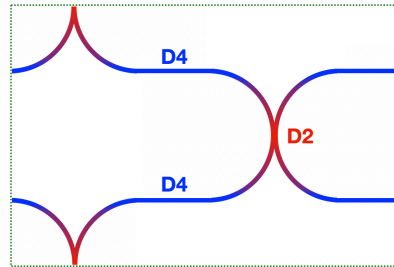
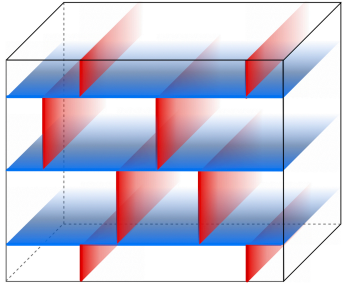
SUPERMAZE

branes pull & merge
16 susy locally !

New Microstate Geometries

$$S = S_{\text{BH}}$$

Effective coupling (g_s)



- Need to build supergravity solution !
- Most generic beast: is 6D sugra enough? or one needs 10D?
- Flat space: supermaze fields **decay exponentially**. Universal ?

What about non-extremal (real-world) BH ?

Extremely hard to build generic microstates

- Coupled 2'nd order PDE's. No susy \Rightarrow **no factorization**
- Numerics, multiple domains, machine learning + Santos
- No susy, no rotation \Rightarrow *factorization !* Bah, Heidmann
- **Schwarzschild microstates geometries !** Bah, Heidmann, Weck
- Cycles with **positive** or **negative** flux (**brane-antibrane**)
- No susy, anti-self-dual fluxes \Rightarrow *factorization !*
- **Running (Kerr-Taub) Bolt** Bena, Giusto, Ruef, Warner 2008
- Non-BPS with rotation. Quadrupole moment **$M_2 > 0$**
Bena, Santos, Pani, Witek; Bena, Lochet
- $M_2^{\text{Kerr}} = -\frac{J^2}{M}$. Normal spinning objects **pancake** $M_2 < 0$
- LIGO data: $+16.0^{+16.7}_{-13.6}$

$$\mathcal{L}_v G_0 = (\mathcal{L}_u G_0) (\partial_z \partial_z G_0) - (\nabla_{\vec{u}} \partial_z G_0) \cdot (\nabla_{\vec{u}} \partial_z G_0)$$

Conclusions

- Give me a lever... to solve Monge-Ampère, and I will give you the e^S horizonless geometries that replace the susy BH horizon.
- So far - big hopes, but **only spikes**.
- “*It is hard for you to kick against the spike*”
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- *KKLT ex nihilo* constraints (talk by Severin)
- Scale separation in AdS solutions is too small for KKLT construction to work
- Swamplanders seem to be right about that one
- What if they are right about inflation as well ?

Waaaaait !

- Haven't Saclay people been arguing since 2009 that *antibranes* \Rightarrow *instabilities*
- Yes ! But unlike de Sitter, non-extremal microstate geometries *should* have instabilities
 - JMaRT (+ bubbles) unstable Cardoso, Dias, Hovdebo, Myers
 - D1-D5: BPS left-moving open string + right-moving open string \Rightarrow emitted closed string
 - Instability time = Hawking Radiation of coherent state !
Chowdhury, Mathur
- Conjectured “swampland” dS instabilities are natural from a BH microscopic perspective.